

Scintillation-based background rejection methods in large scale LAr TPCs

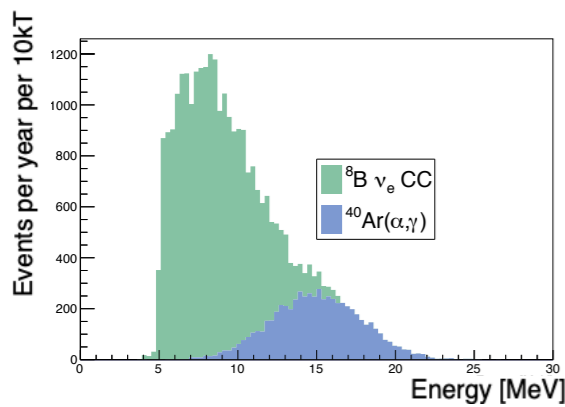
Anyssa Navrer-Agasson, on behalf of the DUNE Collaboration

Motivations

DUNE could deliver world-leading results on solar neutrinos [1]:

- Precision measurement of the ^8B flux
- First observation of the hep flux
- Neutrino-mixing parameters measurements

Background mitigation is essential to achieve high sensitivity.

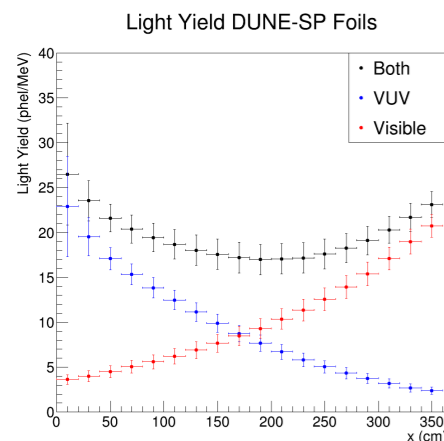
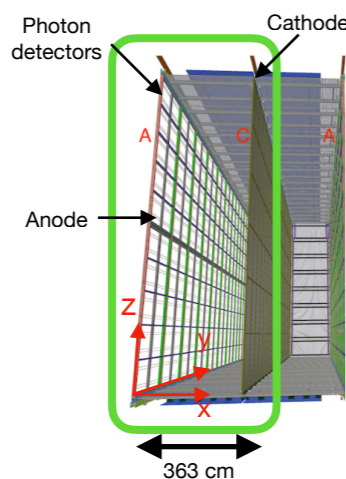


- $^{40}\text{Ar}(\alpha,\gamma)$ reactions have been identified as a challenging background.
- ^{222}Rn is uniformly distributed in LAr.
- α decays from the ^{222}Rn chain can lead to $^{40}\text{Ar}(\alpha,\gamma)$ reactions.
- The produced photons have an energy $E_\gamma \sim 15$ MeV.

Idea: use scintillation light to reduce $^{40}\text{Ar}(\alpha,\gamma)$ backgrounds

Detector simulation

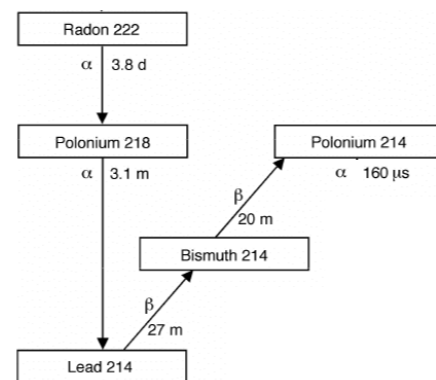
- We consider one section of the DUNE single phase far detector module, between anode and cathode.
- Photon detectors located at the anode plane and 500 V/cm drift field.



- Semi-analytic scintillation model [2].
- Add wavelength shifting reflective foils at the cathode to enhance light collection for events happening far away from the anode.

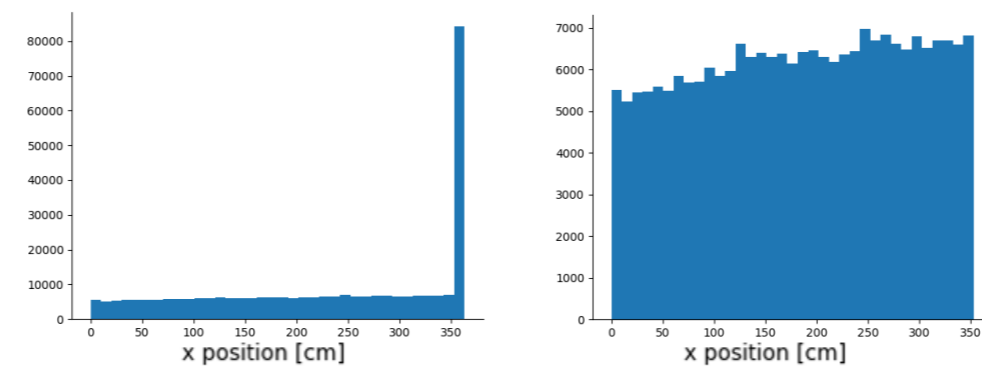
Ion migration model

- Some of the ^{222}Rn daughters are produced as positive ions and migrated to the cathode by the drift field, leading to an accumulation of events near the cathode.
- We can use this effect to discriminate against α -induced backgrounds if we are able to tag events near the cathode.



What we do:

- Start with ^{222}Rn distributed uniformly along x-position with no drift.
- Simulate the decay positions of the different elements of the chain.
- Extract the x distribution of α decays.



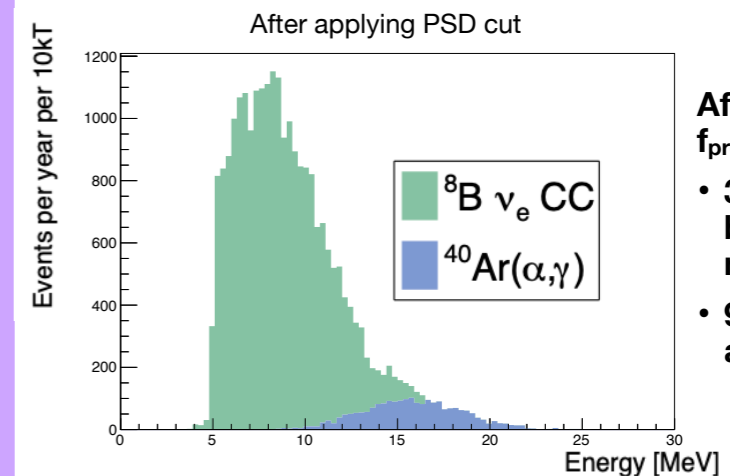
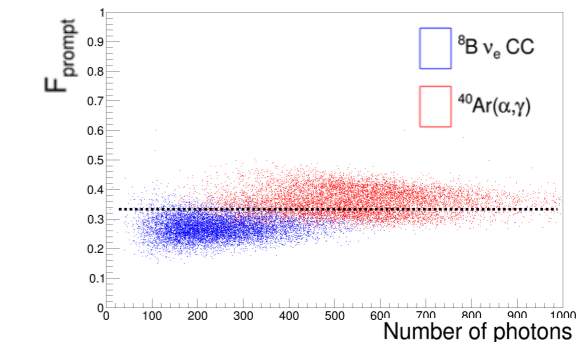
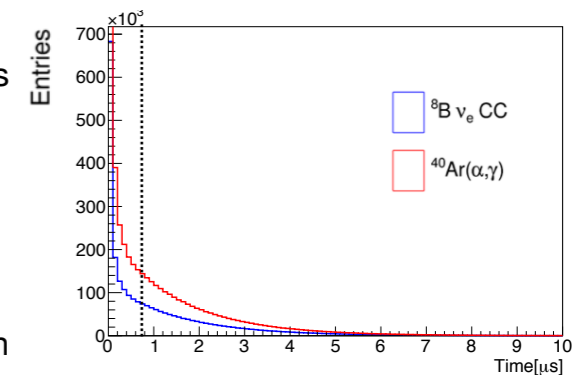
Cut off last bin to show trend

Distance to the cathode	Fraction of α decays
< 30 cm	35%
< 20 cm	32%
< 10 cm	30%

We use the alpha decays position distribution as input to generate $^{40}\text{Ar}(\alpha,\gamma)$ events.

Pulse Shape Discrimination

- ^8B ν_e CC and $^{40}\text{Ar}(\alpha,\gamma)$ interactions have different time profiles.
- We can define a parameter, f_{prompt} , as the fraction of scintillation light received in a certain time window.
- It is then possible to define a cut on f_{prompt} to discriminate between signal and background.



After optimised f_{prompt} cut:

- **34.5% background remaining**
- **97.3% signal acceptance**

References

[1] **DUNE as the Next-Generation Solar Neutrino Experiment**
Francesco Capozzi, Shirley Weishi Li, Guanying Zhu, and John F. Beacom
Phys. Rev. Lett. **123**, 131803

[2] **Predicting transport effects of scintillation light signals in large-scale liquid argon detectors**
Diego Garcia-Gamez, Patrick Green, Andrzej M. Szelc
Eur. Phys. J. C **81** (4) 349 (2021)