

# Electronic versus nuclear recoil discrimination in liquid xenon with PIXeY

Vetri Velan

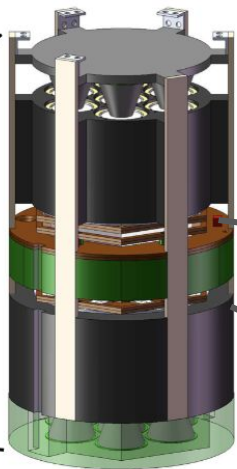
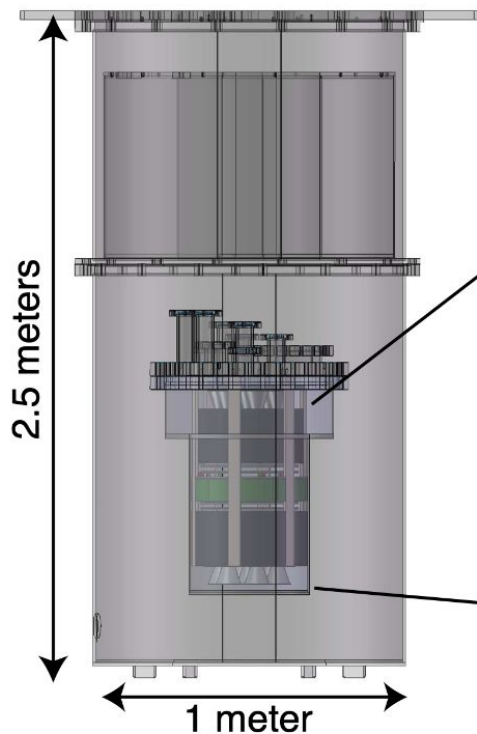
On behalf of the PIXeY Collaboration

Light Detection In Noble Elements (LIDINE)

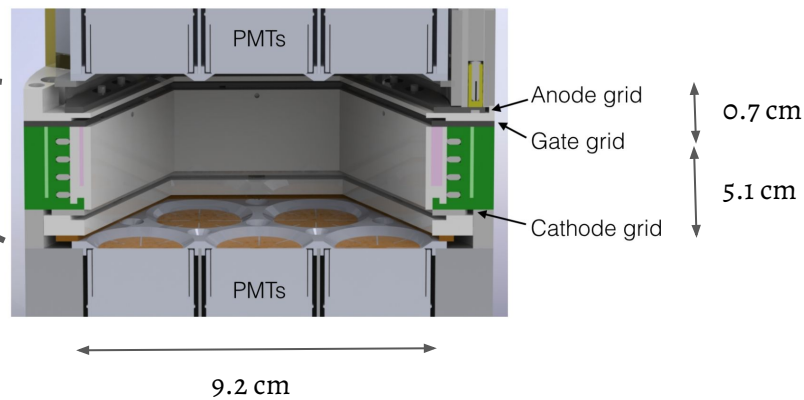
September 15, 2021

**Berkeley**  
UNIVERSITY OF CALIFORNIA

# PIXeY: Particle Identification in Xenon at Yale



- R&D-scale xenon two-phase liquid/gas TPC, operated at Yale in 2014-16
  - 3 kg liquid Xe active volume
- Completed + ongoing analyses on calibrations, extraction efficiency, single electron backgrounds, and **signal/background discrimination**

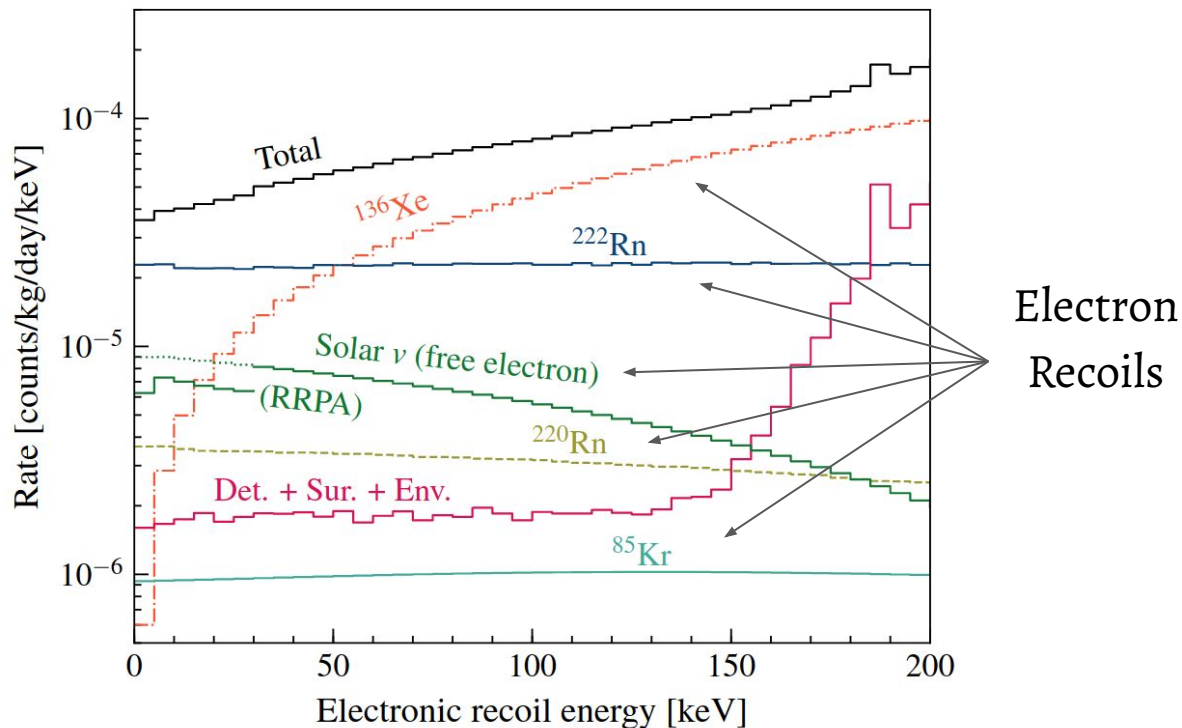


# Backgrounds for Nuclear Recoil Searches

Dominant backgrounds for WIMP dark matter nuclear recoil (NR) searches are electron recoils (ER). Upcoming experiments will face challenging backgrounds that cannot be eliminated by shielding and fiducialization.

## LUX-ZEPLIN Projections

*Phys. Rev. D 101, 052002*

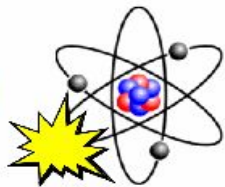


# Electronic Recoil (ER)

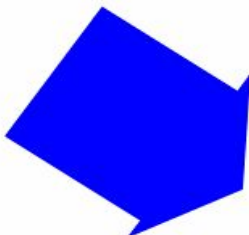
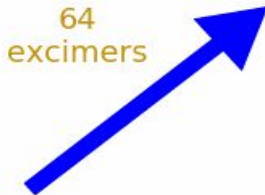
## Energy Deposition

10 keV

200 V/cm



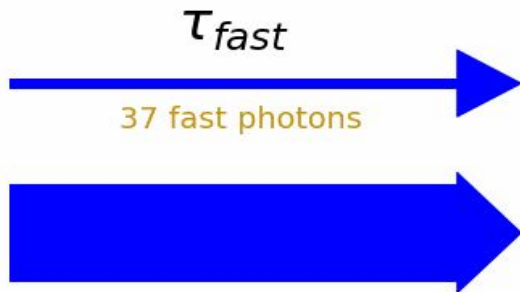
Heat (not observed)



678  
e-ion pairs



286 escaping electrons



**S1**



**S2**

Graphic made with [Xenimation](#)

# ER/NR Discrimination

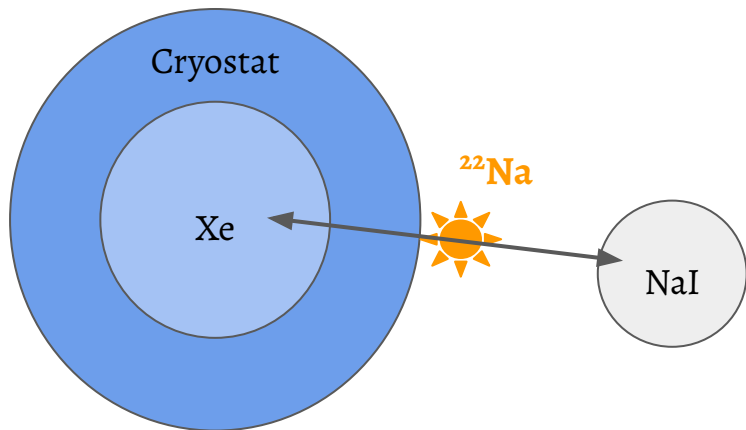
- Discriminate ER from NR by using the ratio of charge-to-light produced by the recoil
  - In practice, discrimination variable is:  $\log_{10}(\mathbf{S2} / \mathbf{S1})$
- PIXeY calibrations for ER and NR allow us to study discrimination as a function of parameters like drift field and energy
- Detector parameters and data parameters for this analysis:
  - Data collected with single-scatter recoils isolated: one S1 pulse, followed by one S2 pulse
  - Additional cuts applied for quality and fiducial volume selection
  - Extraction gas field: **8.0 kV/cm**
  - Drift fields: **117 V/cm** and **213 V/cm**
  - **g1 = 0.110** phe / photon created
  - **g2 = 30.2** phe / escaped electron

# PIXeY Calibrations

Electronic Recoils:  $^{22}\text{Na}$

Gamma source

Tag recoils with an external NaI detector



Nuclear Recoils: **D-D**

2.5 MeV monoenergetic neutrons generated from deuterium fusion

Recoil spectrum  $\sim$  60 GeV WIMP

Detector Effects Calibration:  $^{83\text{m}}\text{Kr}$

Two-step decay, depositing 32.1 and 9.4 keV conversion electrons

Dissolved in liquid xenon; recoil events uniformly fill the detector

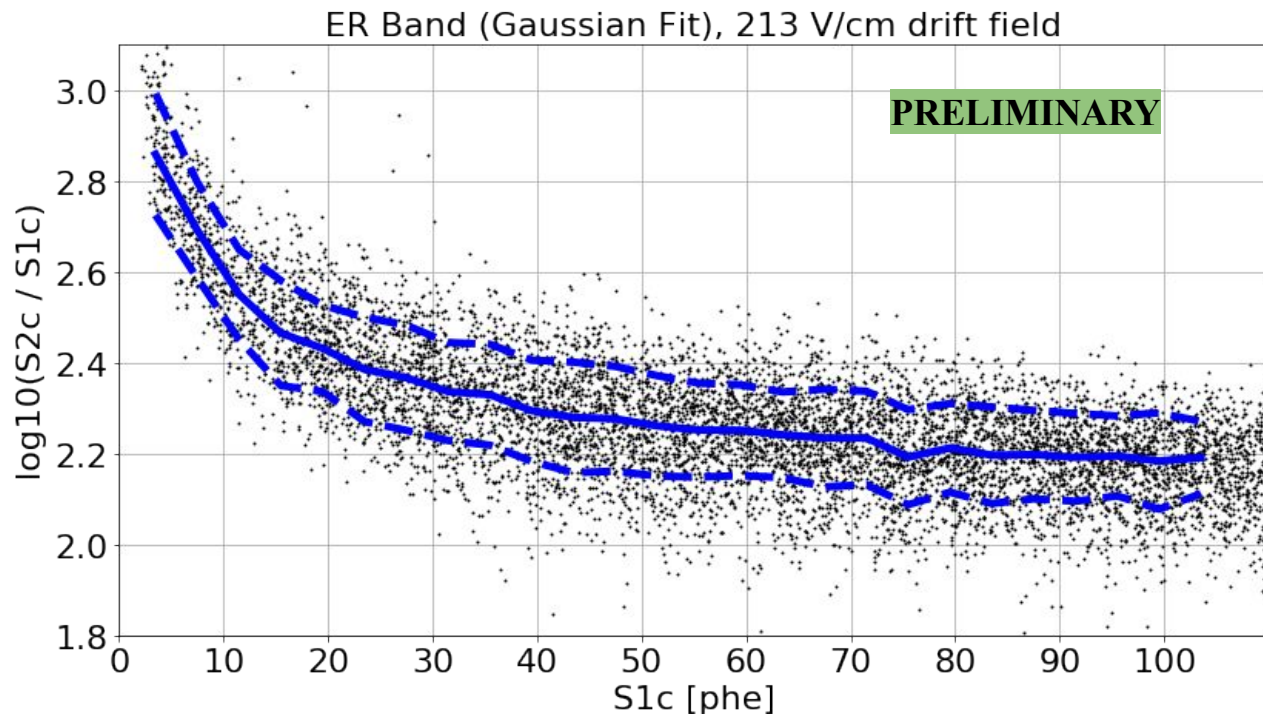
Allows us to identify clean mono-energetic signals and characterize position-dependent and field-dependent phenomena

# Electronic Recoil Band

Events are split into bins of  $S_{1c}$  → the distribution of  $\log_{10}(S_{2c}/S_{1c})$  is fit to a Gaussian

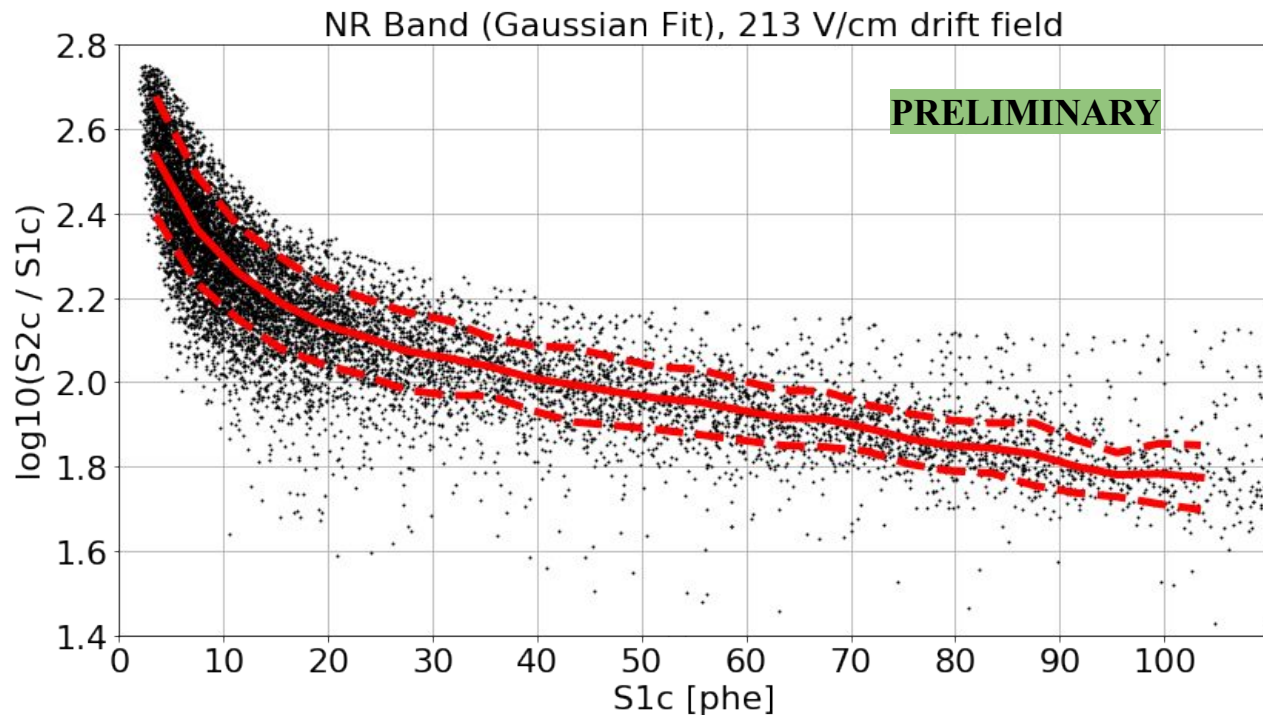
The ER band median and width ( $1\sigma$ ) are extracted

Skewness ignored for now



# Nuclear Recoil Band

The NR band median and width are extracted the same way





# Leakage and Discrimination

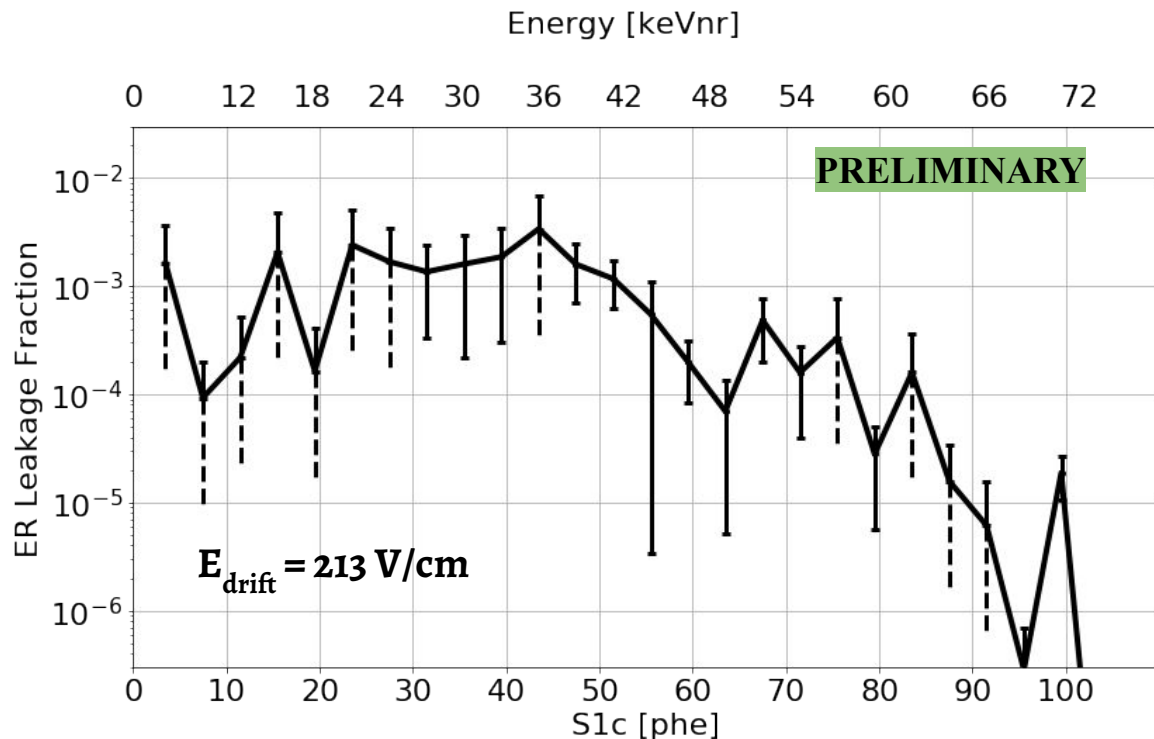
Extrapolate the ER band Gaussian fit below the NR band to calculate leakage fraction

Leakage improves dramatically with energy, as seen by others (e.g. LUX)

$\leq 10^{-6}$  leakage at  $S1 = 100$  phe

Leakage over the WIMP search range (0-80 phd) is  $1 \times 10^{-3}$ , substantially better than the LZ requirement of  $5 \times 10^{-3}$

(LZ requirement in [arxiv:1509.02910](https://arxiv.org/abs/1509.02910))



# LUX Comparison

PIXeY:  $g_1 = 0.110$ ,  
drift field = 213 V/cm

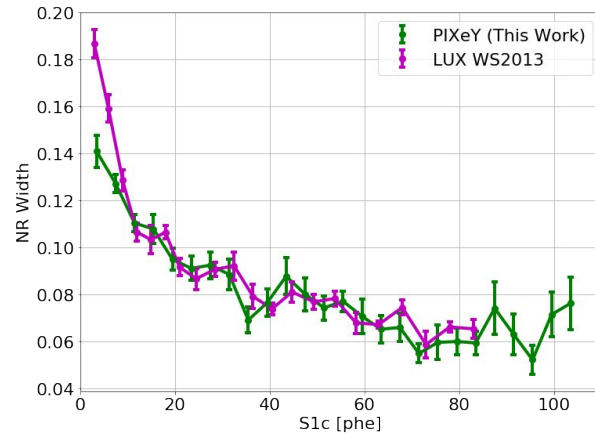
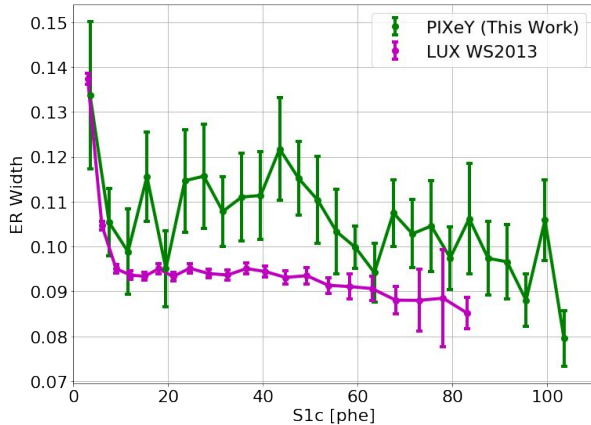
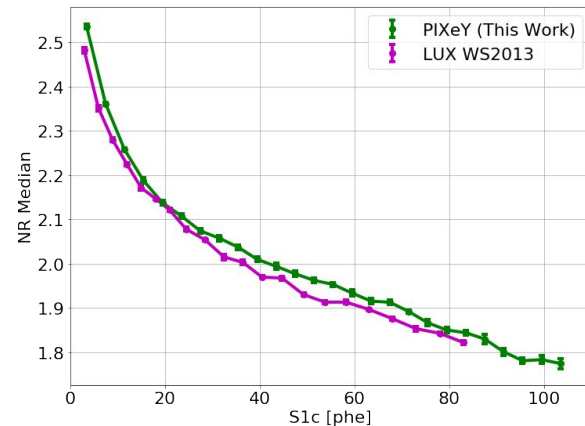
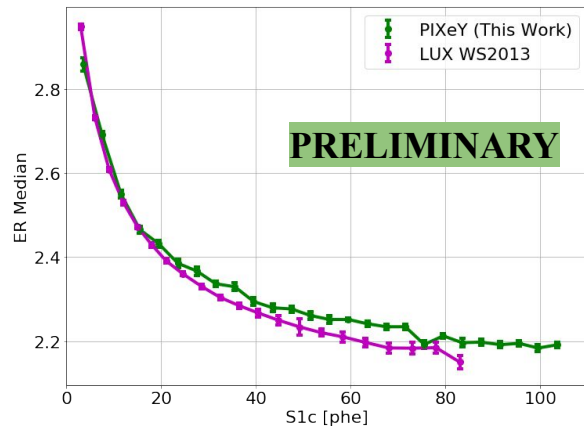
LUX 2013:  $g_1 = 0.117$ ,  
drift field = 180 V/cm

Lower band median and width in  
LUX make sense, because of higher  
 $g_1$  and lower field

LUX “ER width” is  $\sigma_-$ , the width of  
the downwards fluctuations. By  
definition:  $\sigma_- < \sigma$ . Thus using  $\sigma_-$  with  
skewness will give better leakage.

(LUX data from Phys. Rev. D 102, 112002 (2020))

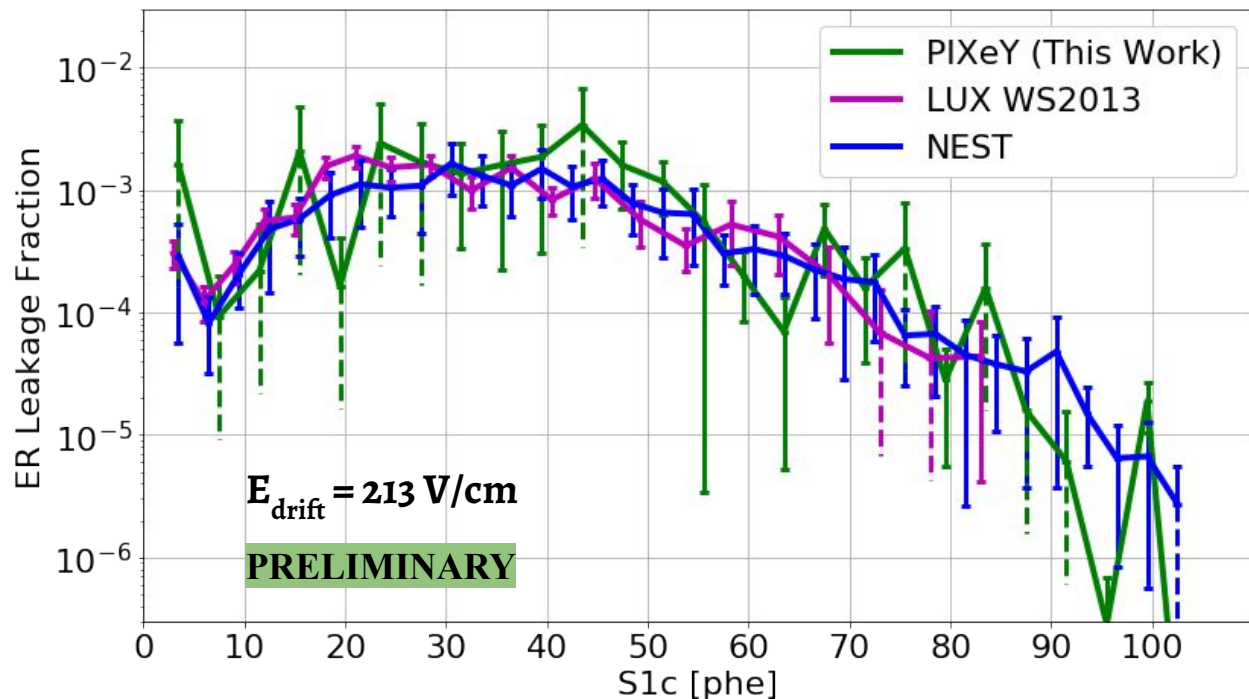
$E_{\text{drift}} = 213 \text{ V/cm}$



# Comparison to LUX Data and NEST Simulation

Simulated PIXeY with NEST, using correct detector parameters for  $g_1$ ,  $g_2$ , drift and extraction fields, etc.

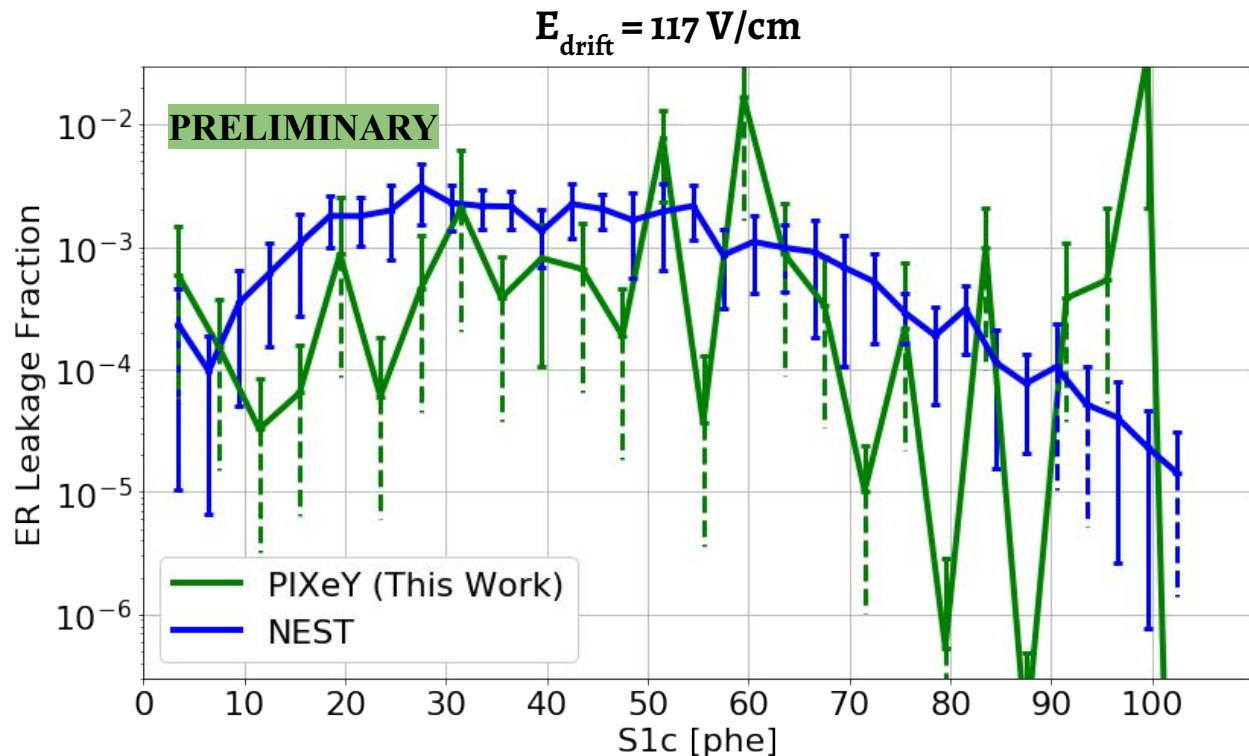
PIXeY leakage matches LUX and NEST remarkably well!



# Varying Drift Electric Field

Analysis repeated with a drift field of 117 V/cm

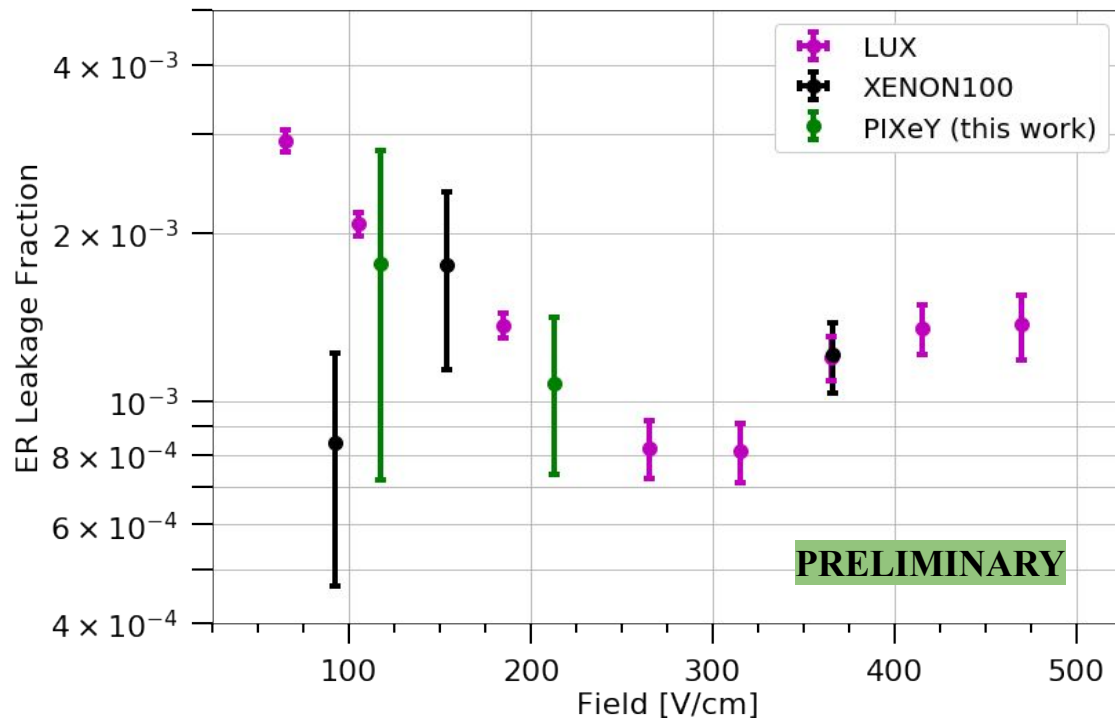
Matches NEST again, albeit with larger statistical uncertainties



# Varying Drift Electric Field

LUX data from Phys. Rev. D 102, 112002 (2020)

XENON100 data from Phys. Rev. D 97, 092007 (2018)



# Summary and Conclusions

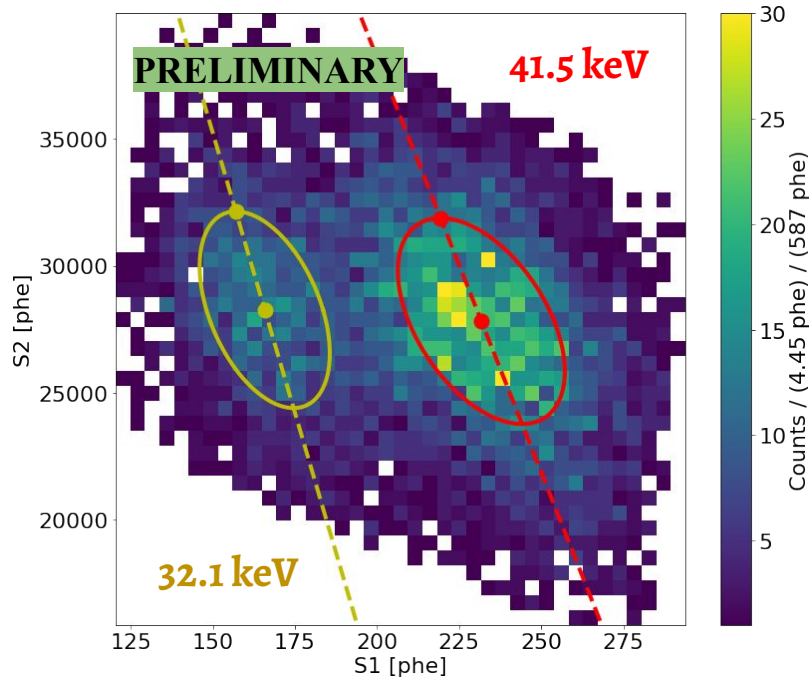
- The PIXeY experiment demonstrates robust ER/NR discrimination at high energies, achieving a factor of  $\sim 10^6$  ER rejection at  $S1 = 100$  phe ( $E = 70$  keVnr)
- Results are consistent with past experiments, including LUX and XENON100
- Future discrimination work will focus on ER skewness and further variation of drift field, up to the  $\sim$ kV/cm region

**Thank you!**  
**Any questions?**

# Backup

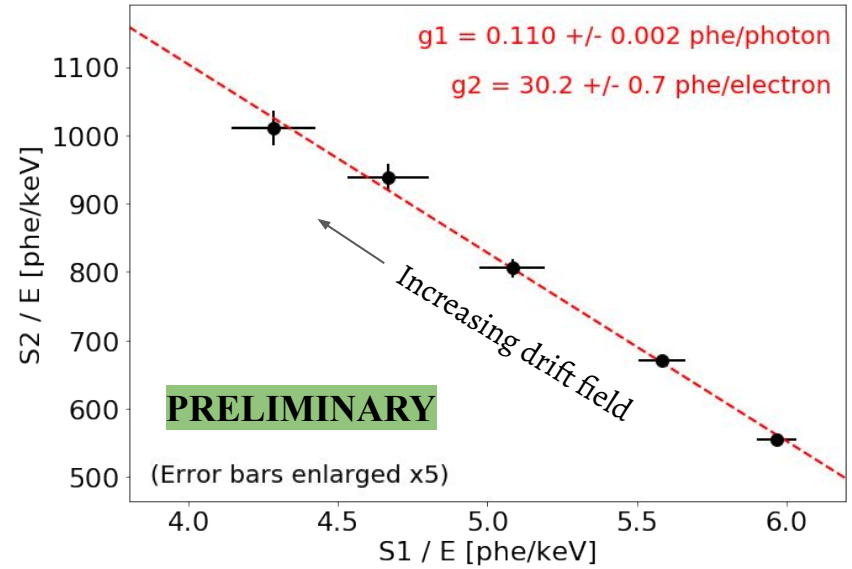


# PIXeY Detector Parameters



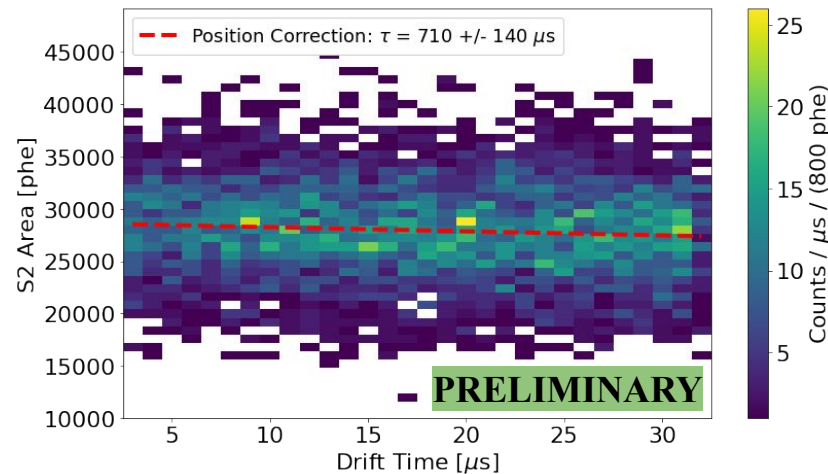
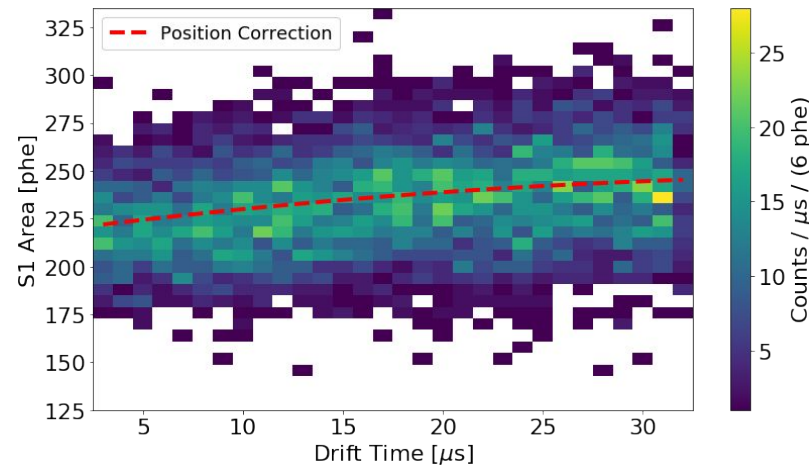
Calculate  $g_1$  and  $g_2$  using the “Doke” technique

- These represent the average signal size for each photon and electron leaving the recoil site



# Position Corrections

- Pulse areas are dependent on the z-position of the event; S1 sizes grow with drift time, while S2 sizes fall with drift time
- S1 sizes grow with drift time (i.e. shrink with height) due to total internal reflection at the liquid-gas interface
- S2 sizes shrink with drift time (i.e. grow with height) due to ionized electrons capturing on electronegative impurities

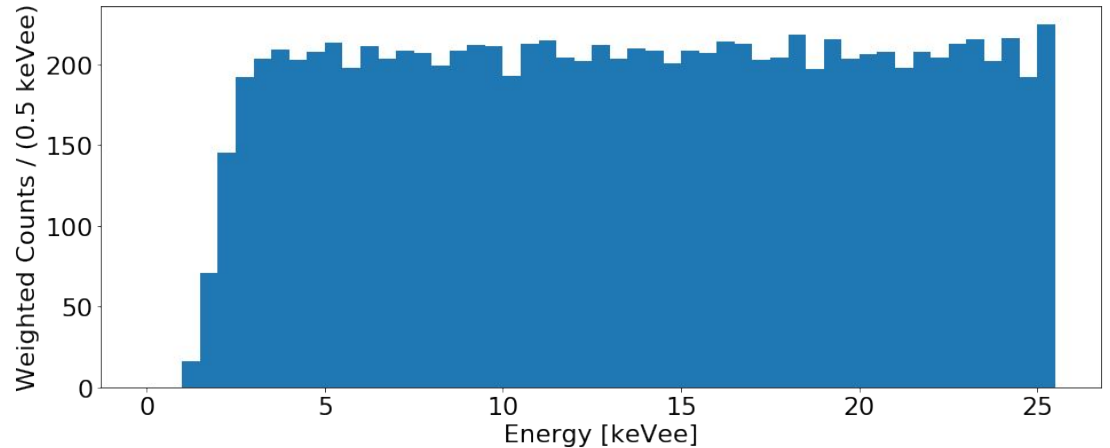
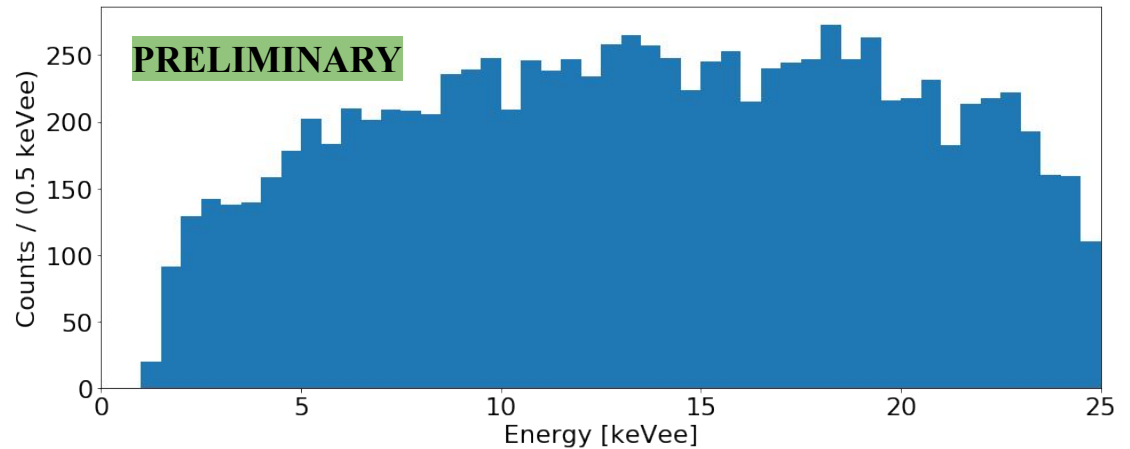


# Energy Weighting

Electronic recoil events are weighted to mimic a flat spectrum in combined energy

Dark matter ER backgrounds from  $^{136}\text{Xe}$  double beta decay, Rn/Kr internal contaminants, and solar neutrinos are roughly flat in energy

$$E = W (S1c / g1 + S2c / g2)$$



# Noble Element Simulation Technique (NEST)

- Inter-collaboration collaboration of liquid xenon and argon physicists who construct models of energy deposition in a noble detector, as well as code to implement simulations
  - Members from LUX, LZ, XENON, (n)EXO, RED100, COHERENT, DUNE, ICARUS, MicroBooNE, SBN
- Models are based on physical principles and world averages of existing data
- PIXeY simulation:
  - Template detector file, which represents LUX WS2013
  - Detector parameters like  $g_1$ ,  $g_2$ , drift field, extraction efficiency were adjusted to their correct values
  - Simulated a flat ER energy distribution and DD neutrons → calculated ER + NR bands and leakage fraction
- Code and more information:
  - <http://nest.physics.ucdavis.edu/>
  - <https://github.com/NESTCollaboration>

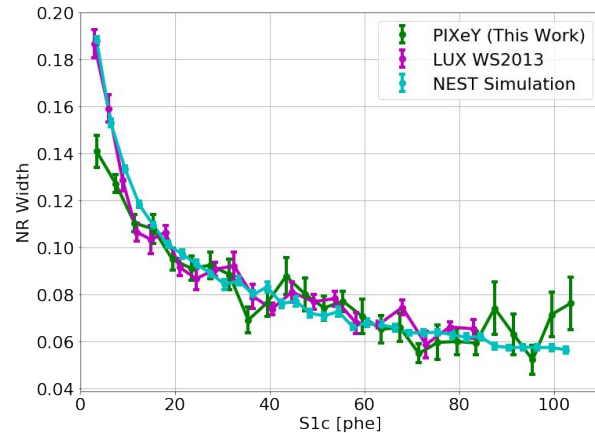
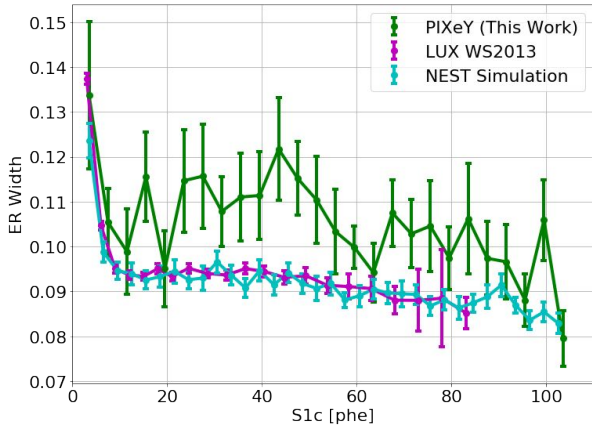
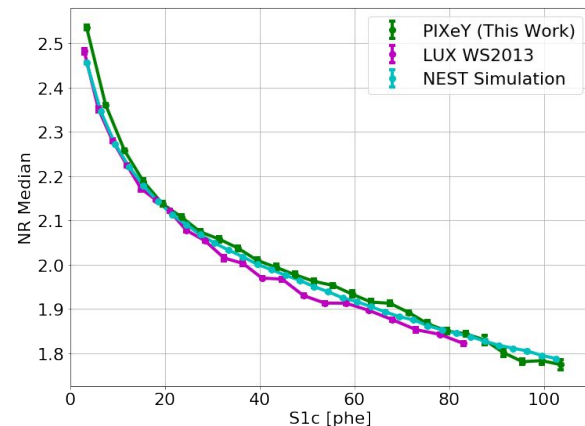
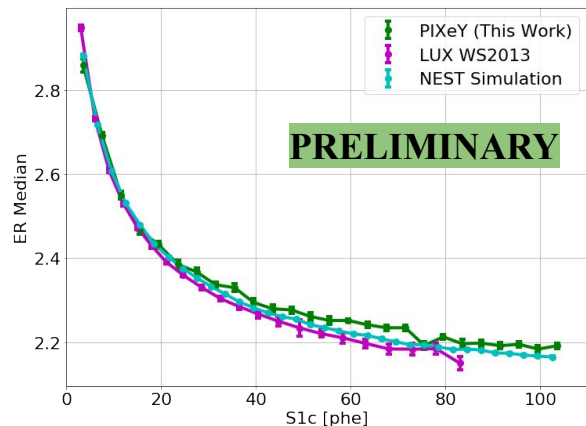
# NEST Simulation

Used NEST to simulate the PIXeY experiment, with proper  $g_1$ ,  $g_2$ , drift field, and other detector parameters

Matches pretty well, although there are some discrepancies

NEST matches LUX 2013 very well, although the detector parameters are slightly different

$$E_{\text{drift}} = 213 \text{ V/cm}$$

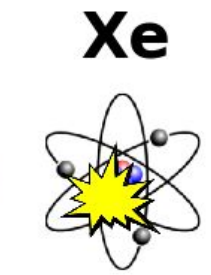


# Nuclear Recoil (NR)

## Energy Deposition

10 keV

200 V/cm



Heat (not observed)

62  
excimers



77  
e-ion pairs



58 escaping electrons



$\tau_{fast}$

18 fast photons

**S1**

$\tau_{slow}$

62 slow photons

18 recombining electrons

**Recombination**

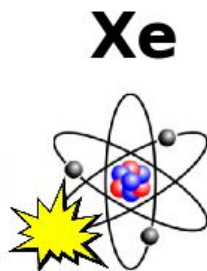
Graphic made with [Xenimation](#)

# Electronic Recoil (ER)

## Energy Deposition

10 keV

200 V/cm



64  
excimers



$\tau_{fast}$

37 fast photons

**S1**

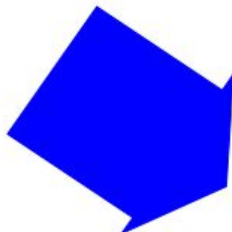


$\tau_{slow}$

419 slow photons

392 recombining electrons

**Recombination**



678  
e-ion pairs



**S2**

286 escaping electrons

Heat (not observed)

Graphic made with [Xenimation](#)