

# Sensitivity of the nEXO neutrinoless double beta decay experiment

(Based on [arXiv:2106.16243](https://arxiv.org/abs/2106.16243))

Brian Lenardo

Light Detection In Noble Elements (LIDINE)

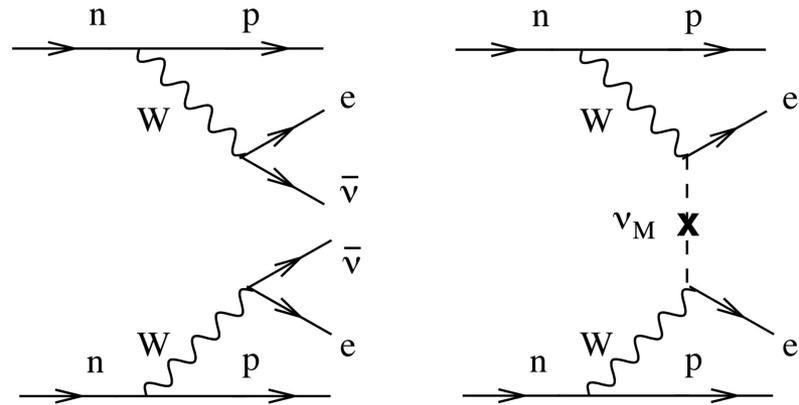
September 14, 2021



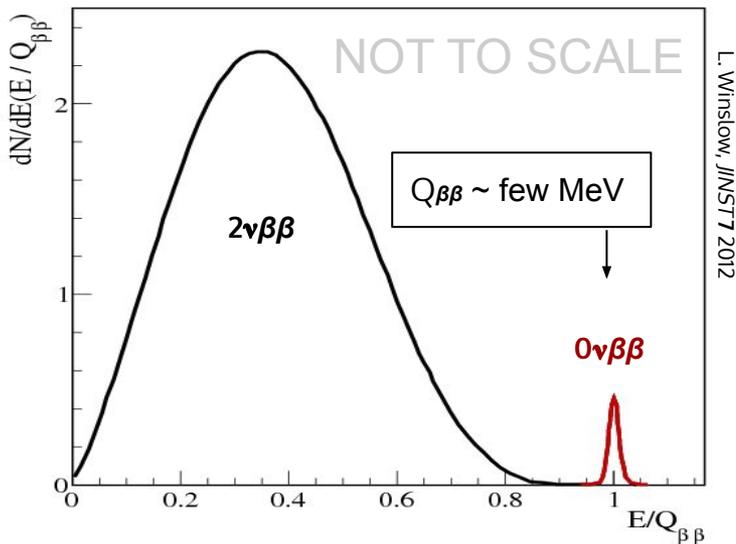
# Searching for new physics via $0\nu\beta\beta$

$0\nu\beta\beta$  decay is a sensitive probe of physics beyond the standard model

- Lepton number violation
- Majorana nature of neutrino
- May explain neutrino mass scale (see-saw mechanism), matter/antimatter asymmetry (leptogenesis)



Avignone, Elliott, & Engel  
*Rev. Mod. Phys.* **80** (2008)

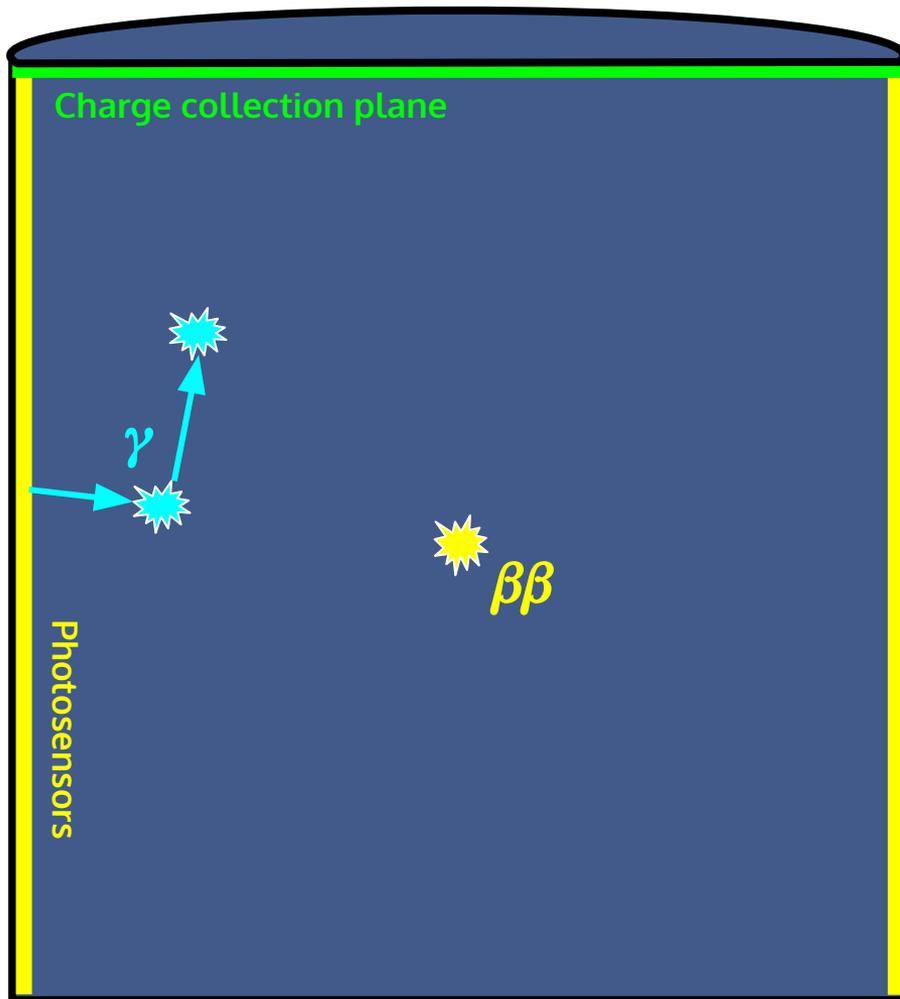


**Current  $T_{1/2}$  limits:  $>10^{26}$  years!**

**Requirements for next-gen  $0\nu\beta\beta$  experiments:**

- A LOT of the isotope of interest ( $>10^{27}$  atoms)
- Low backgrounds in MeV range (low radioactivity)
- Signal/background discrimination
- Good energy resolution

# Searching for $0\nu\beta\beta$ in a liquid Xe TPC



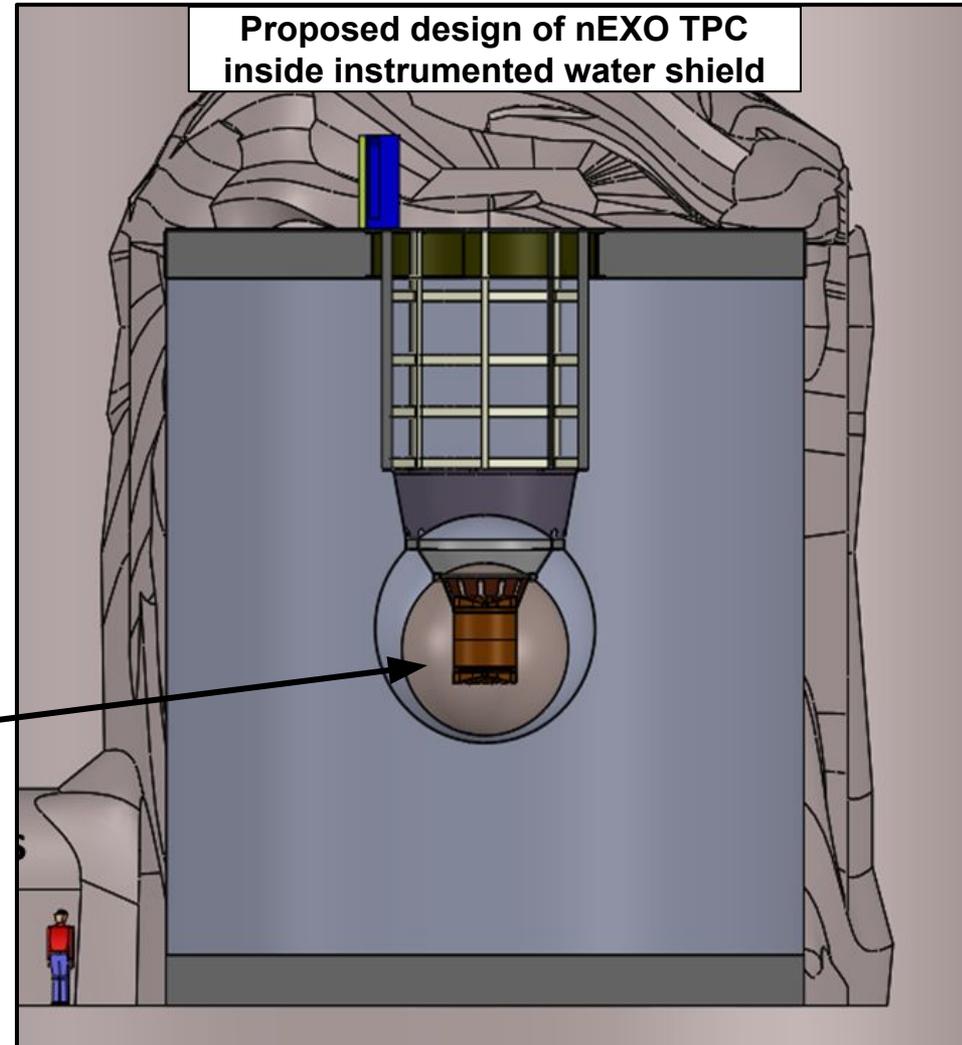
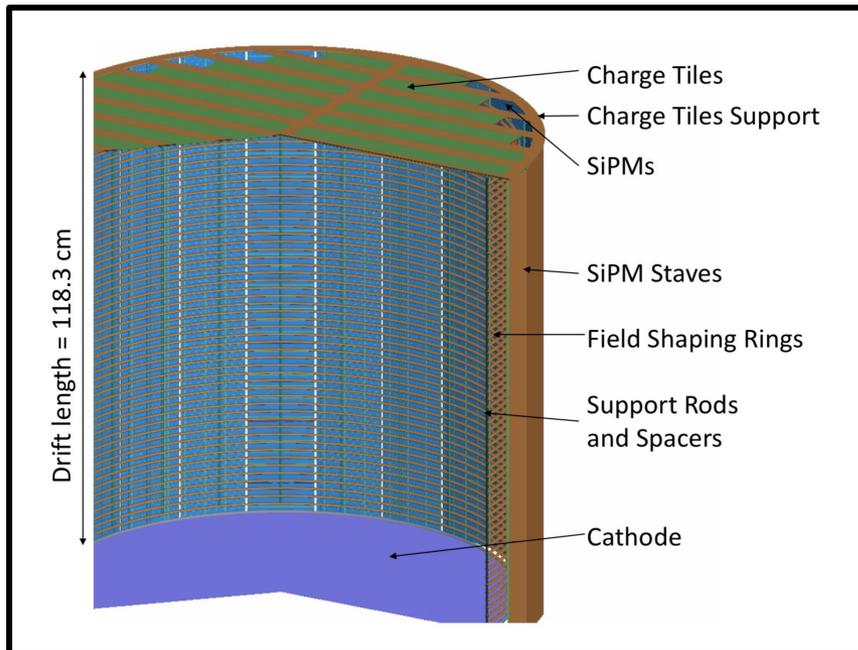
## Addressing the challenges in $0\nu\beta\beta$ searches:

- Large samples of  $\beta\beta$  isotope ( $^{136}\text{Xe}$ )
  - Ton-scale LXe TPCs are already operating
  - Enrichment is straightforward
- Low, well-characterized backgrounds in MeV range (low radioactivity)
  - Very low intrinsic backgrounds in Xe
  - Strong self-shielding
- Good energy resolution
  - Combining charge and light can reach <1%
- Signal/background discrimination
  - Powerful position reconstruction and multi-site rejection

# The next-gen Enriched Xenon Observatory (nEXO)

Search for  $0\nu\beta\beta$  decay of  $^{136}\text{Xe}$  ( $Q_{\beta\beta} = 2.457$  MeV)

- Single-phase liquid xenon TPC
- Five tonne target volume, 90% enriched in  $^{136}\text{Xe}$
- ~2000m overburden
  - To be located in the SNOLAB cryopit
- Active water Cherenkov veto



# Projecting the sensitivity of nEXO

## Procedure:

1. Build a background model
2. Simulate the detector response
3. Analyze (simulated) data and perform statistical analysis

# Projecting the sensitivity of nEXO

## Procedure:

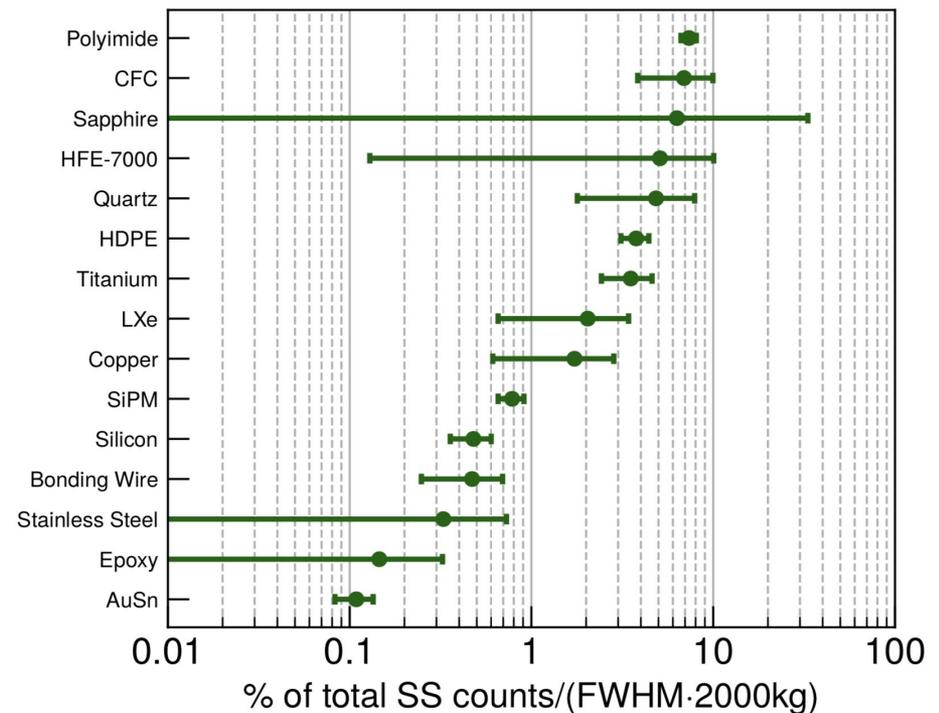
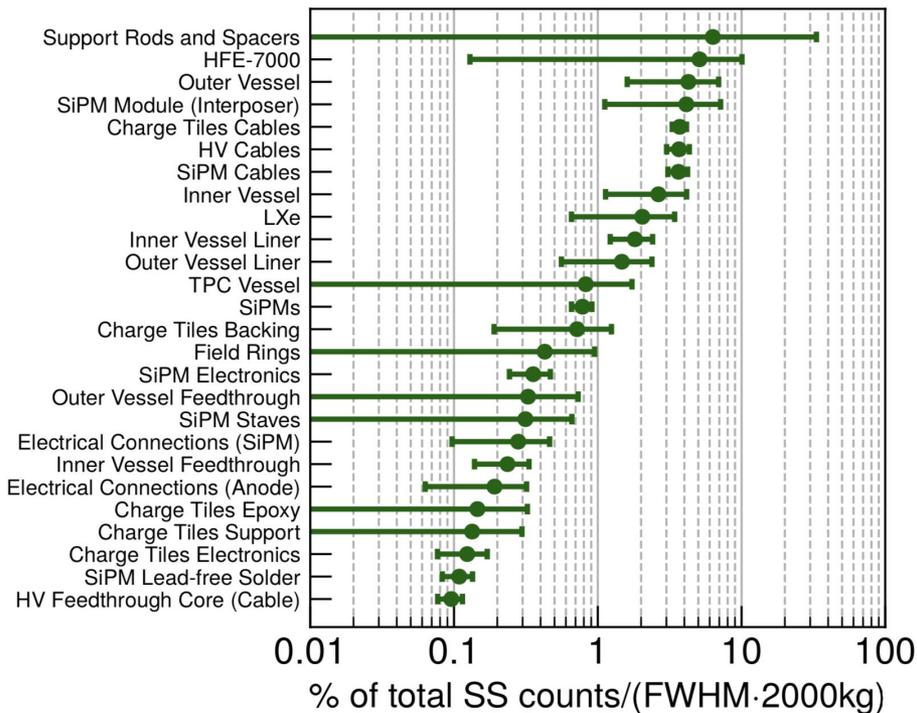
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# 1. Building a background model: materials

Extensive cleanliness campaign, following success of EXO-200 (e.g. D. Leonard, *NIMA* 591 (2008) )

Every material in existing design has been screened for radiopurity

→ nEXO's material background assumptions are data-driven

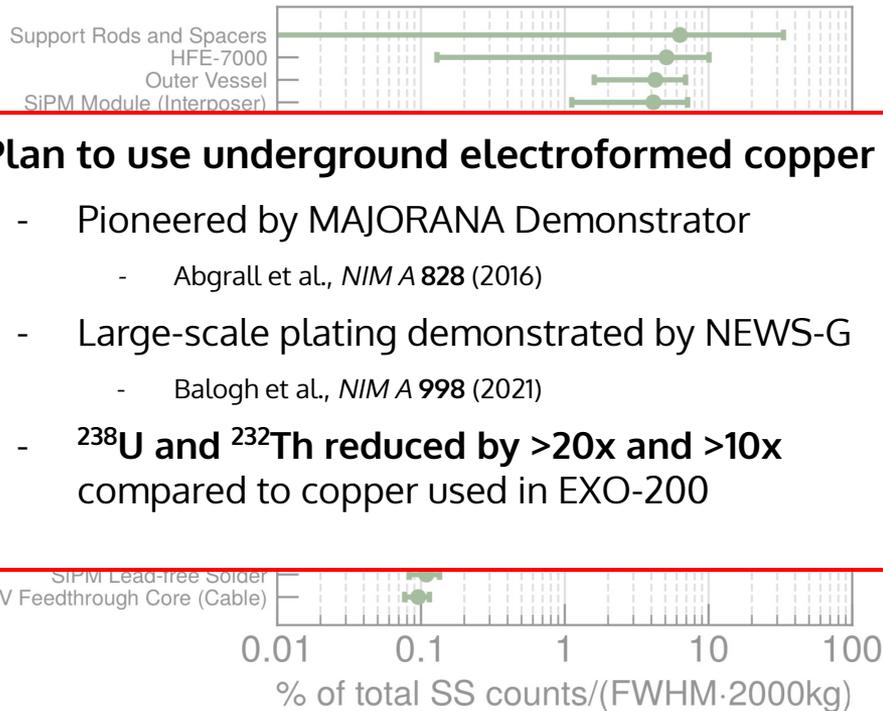


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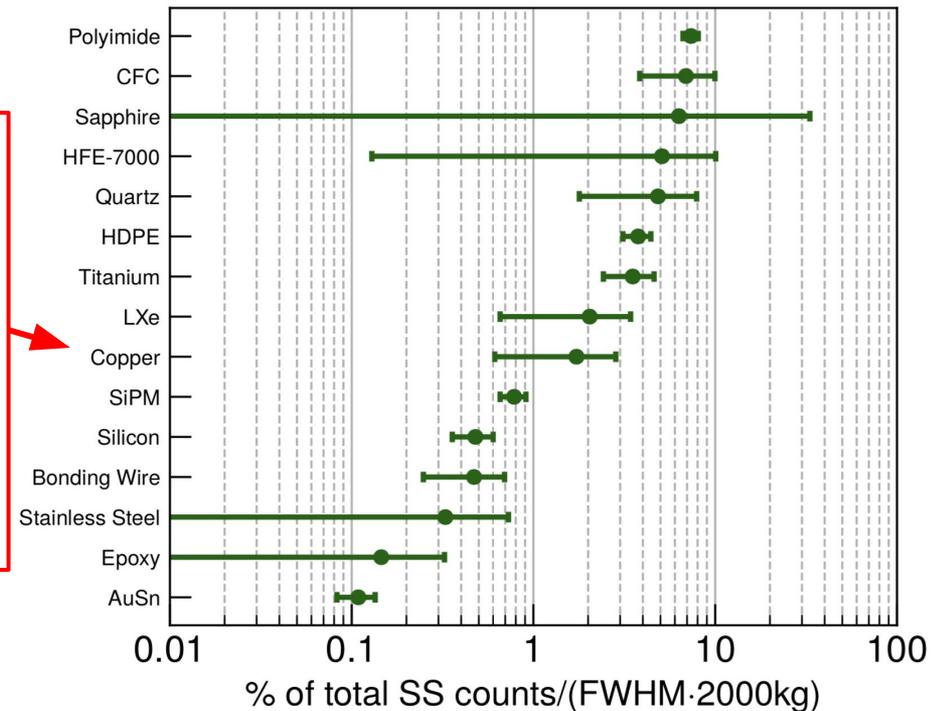
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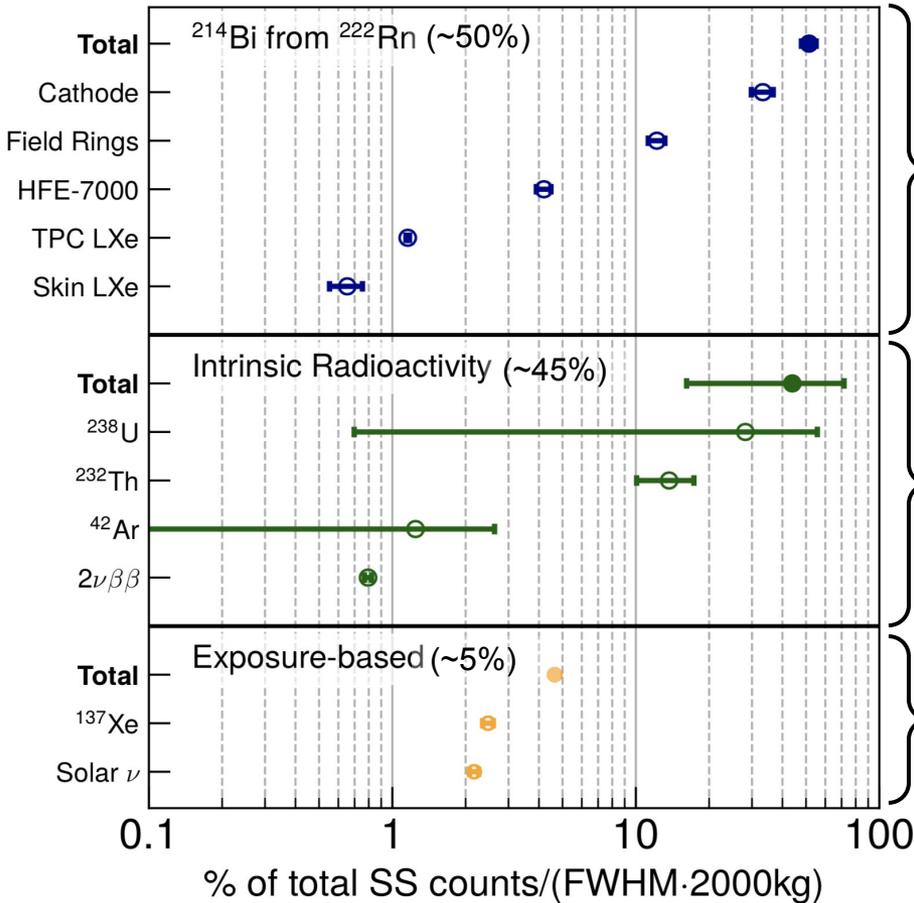
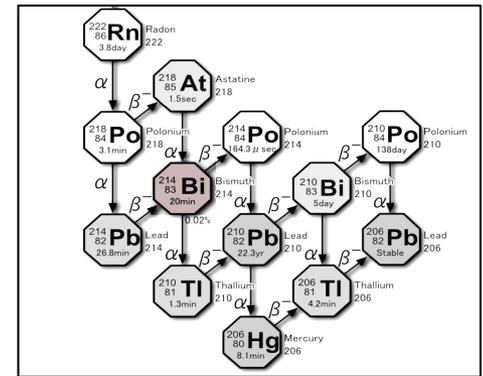


## Plan to use underground electroformed copper

- Pioneered by MAJORANA Demonstrator
  - Abgrall et al., *NIM A* 828 (2016)
- Large-scale plating demonstrated by NEWS-G
  - Balogh et al., *NIM A* 998 (2021)
- $^{238}\text{U}$  and  $^{232}\text{Th}$  reduced by >20x and >10x compared to copper used in EXO-200



# 1. Building full background model



## Primary bkg from $^{222}\text{Rn}$ emanated into the Xe system

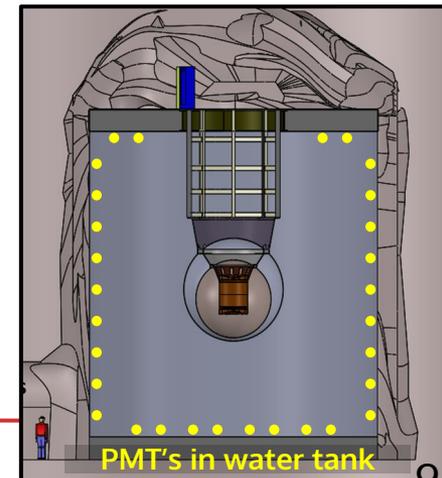
- Assume steady-state population of 600 atoms of  $^{222}\text{Rn}$  in the xenon (extrapolation from EXO-200)
- Primary issue is daughters attached to detector surfaces

## Detector materials

- Previous slide

## Backgrounds unaffected by detector construction

- We conservatively assume **70% rejection of  $^{137}\text{Xe}$  bkg** via active muon veto, based on simulations



# Projecting the sensitivity of nEXO

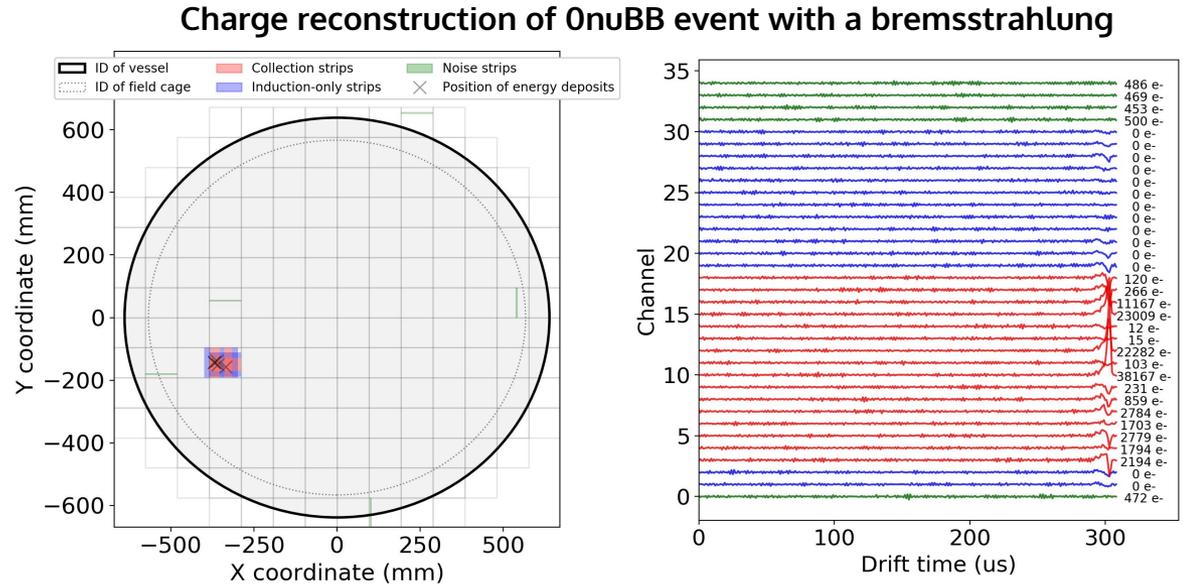
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# Recent work on readout simulations

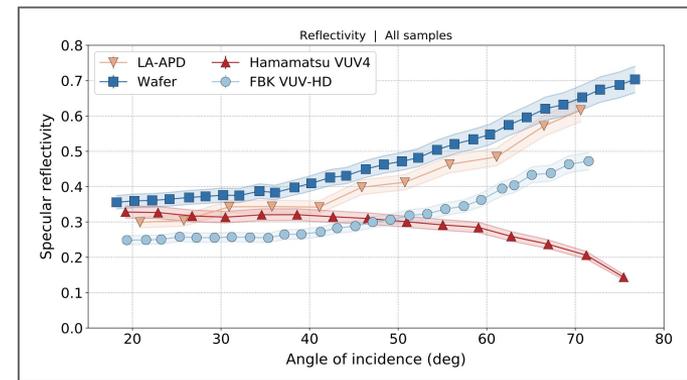
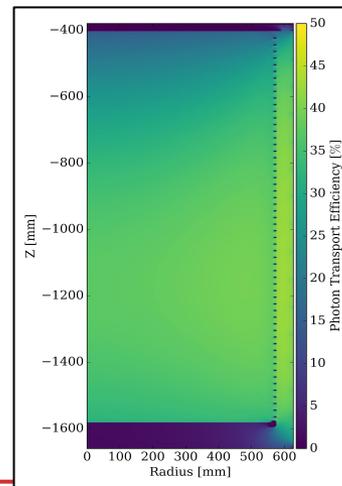
## Charge readout modeled at the waveform level

- Full noise simulation of ASIC readout and charge propagation through TPC
- Induction and noise signals generated to mimic real data
  - Z. Li et al. (NEXO) *JINST* **14** (2019)



## Light collection modeling improved with new data and software

- High-stats, fine-grained simulation using GPU-based Chroma software
- New measurements of SiPM optical properties and performance
  - P. Nakarmi et al. (NEXO) *JINST* **15** (2020)
  - G. Gallina et al. (NEXO) *NIMA* **940** (2019)
  - A. Jamil et al. (NEXO) *IEEE TNS* **65** (2018)
  - P. Lv et al. (NEXO) *IEEE TNS* **67** (2020)
  - M. Wagenpfeil et al. (NEXO) *JINST* **16** (2021)



M. Wagenpfeil, *JINST* (2021) arXiv:2104.07997

# Projecting the sensitivity of nEXO

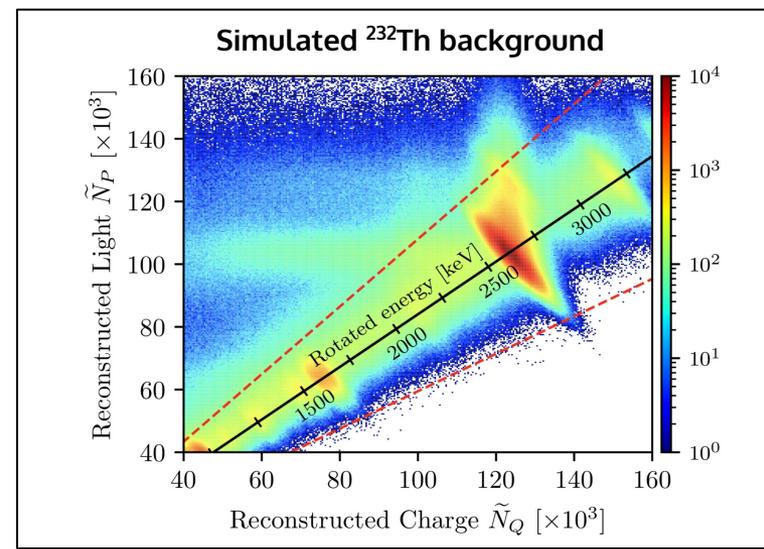
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# Data analysis in 3D

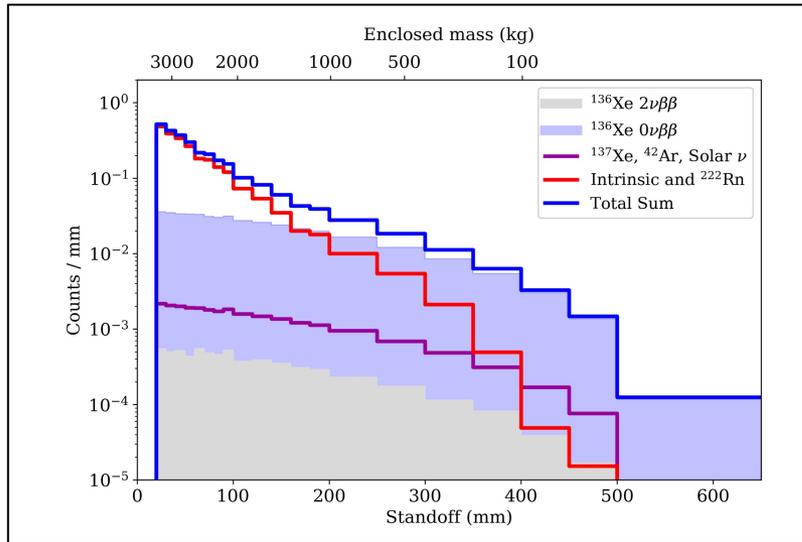
## Event energy:

- Sum of reconstructed light and charge



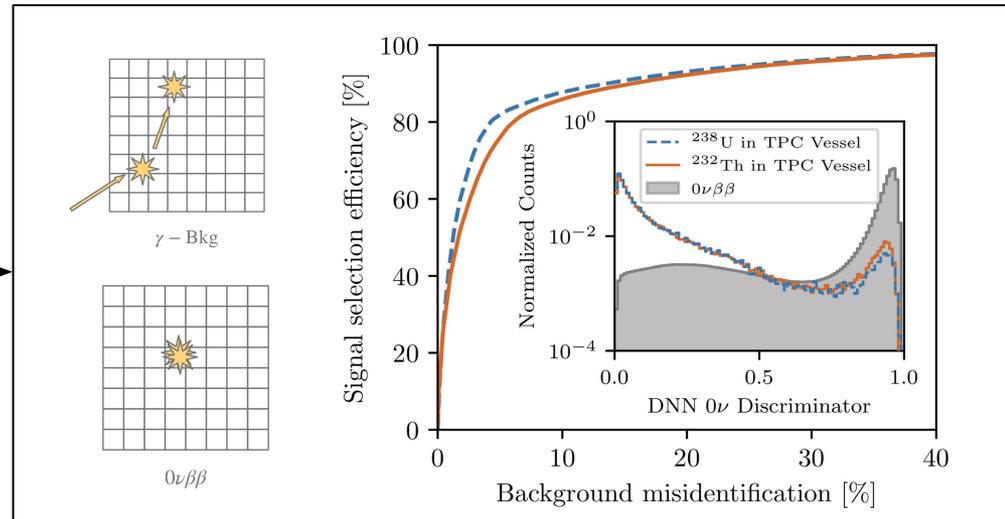
## Standoff distance:

- Defined as minimum distance between an event vertex and a detector surface



## Deep neural network (DNN):

- Utilize imaging capabilities of TPC to separate **single-site** (signal-like) from **multi-site** (background-like)



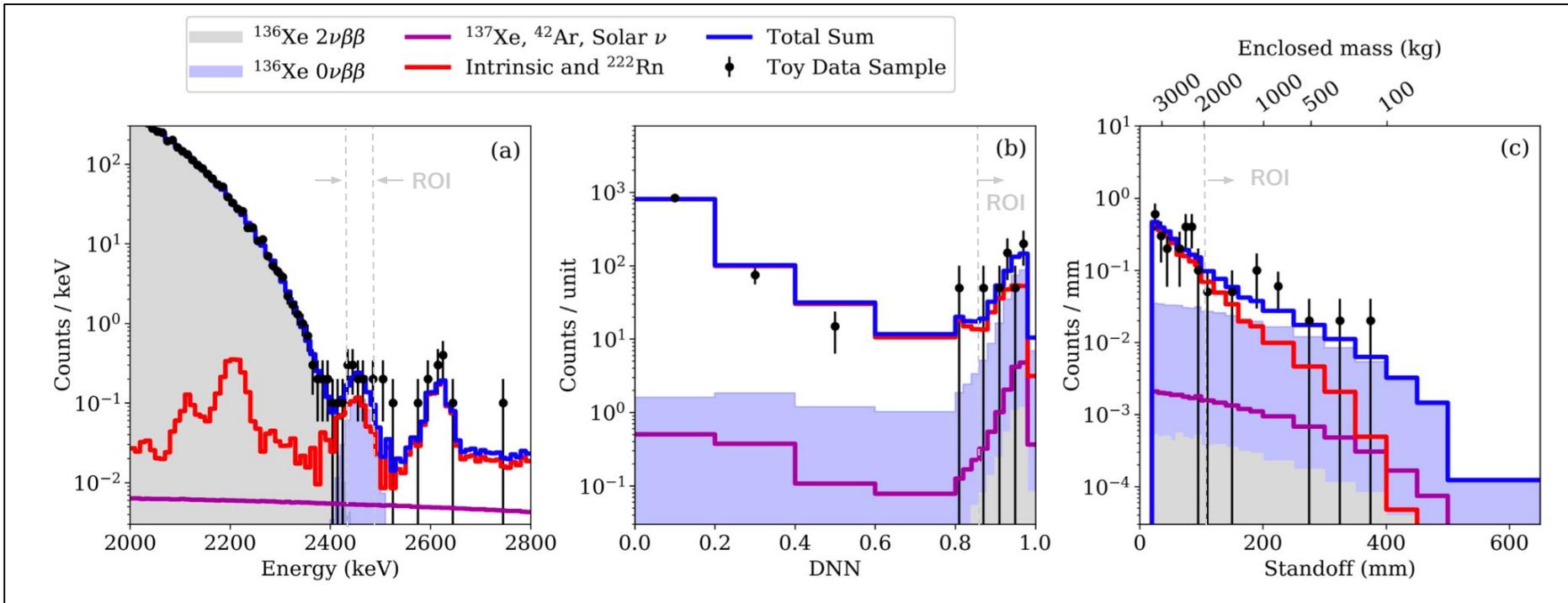
# nEXO's data inside the "ROI"

Illustration of a simulated dataset with a  $3\sigma$  discovery ( $T_{1/2} = 0.74 \times 10^{28}$  yrs)

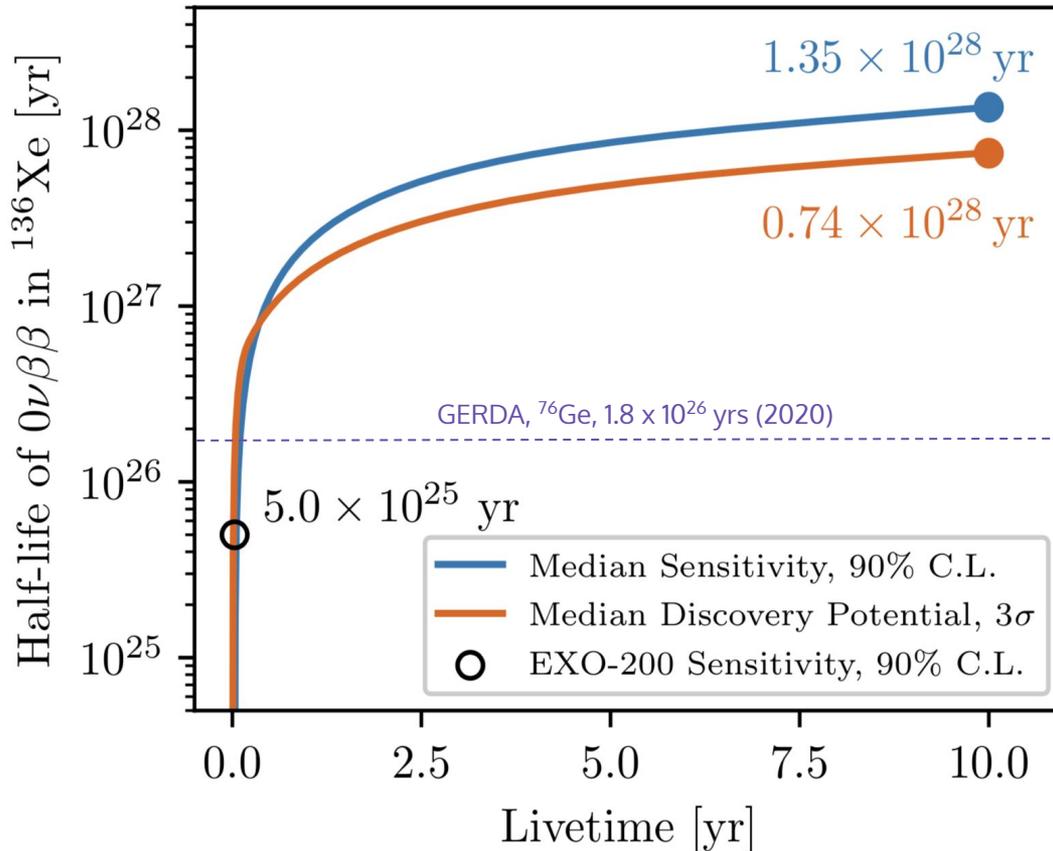
Three-dimensional analysis allows robust discovery of a signal

nEXO's ROI is defined as:

- Energy within  $0\nu\beta\beta$  FWHM
- Innermost 2 tonnes of liquid Xe (cut on standoff)
- DNN > 0.85 ("single-site" events)



# Sensitivity and discovery potential for $0\nu\beta\beta$



## Three-dimensional profile likelihood ratio analysis

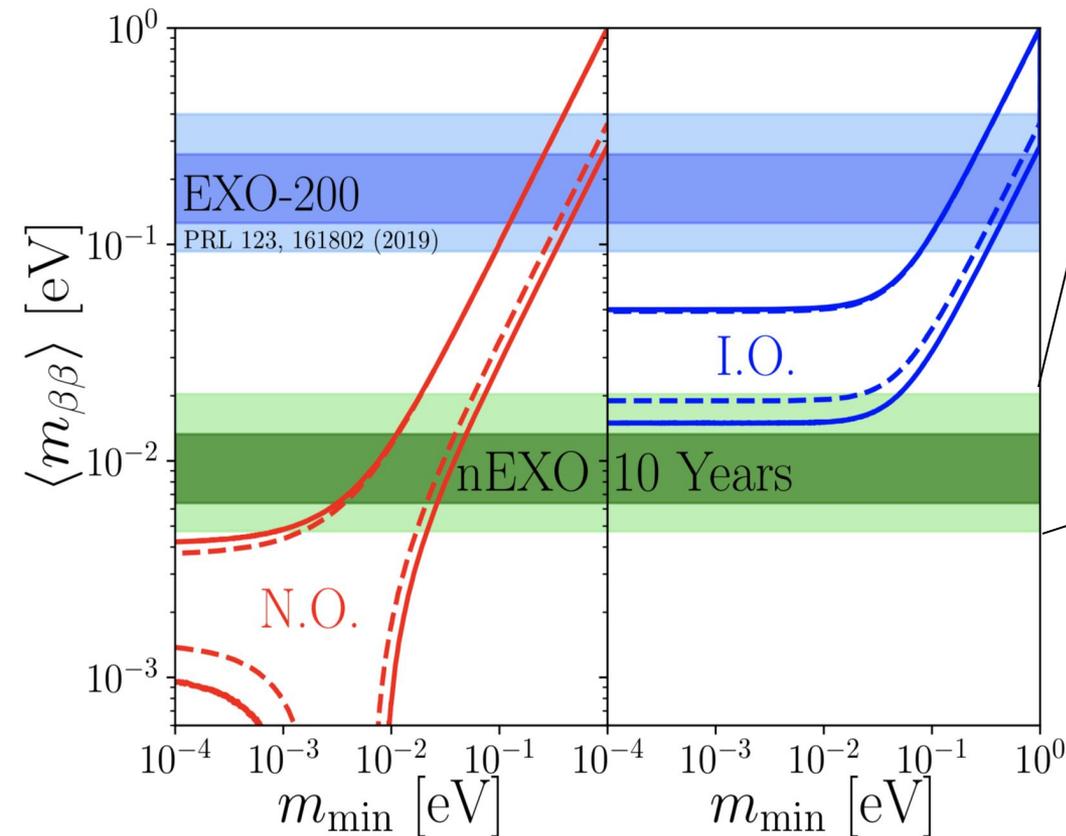
- Optimizes discrimination between signal and background
- Robust against previously unknown gamma backgrounds

## Sensitivities reported are median values over 5,000 toy experiments

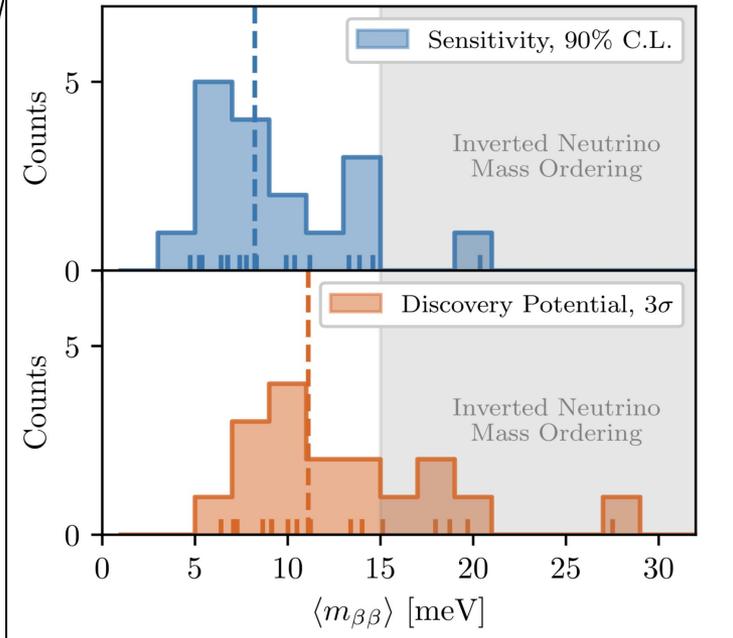
- Can *exclude*  $0\nu\beta\beta$  halflife of  **$1.35 \times 10^{28}$  yrs at 90% CL**
- Can *discover*  $0\nu\beta\beta$  halflife of  **$0.74 \times 10^{28}$  yrs at  $3\sigma$  significance**

# Sensitivity to new physics

$$\left[ T_{1/2}^{0\nu} \right]^{-1} = \frac{\langle m_{\beta\beta} \rangle^2}{m_e^2} G^{0\nu} \mathcal{M}^{0\nu}$$



Distribution over different matrix element calculations



Sensitivity to Majorana neutrino mass:

$$m_{\beta\beta} \approx 4.7 - 20.3 \text{ meV}$$

Completely explores the inverted mass ordering in almost all cases

# Conclusions

nEXO will employ a five-tonne liquid xenon time projection chamber to search for neutrinoless double beta decay

- Leverages well-established capabilities of xenon TPCs

Projected sensitivity of  $T_{1/2} > 1.35 \times 10^{28}$  yrs at 90% CL, nearly two orders of magnitude beyond current limits

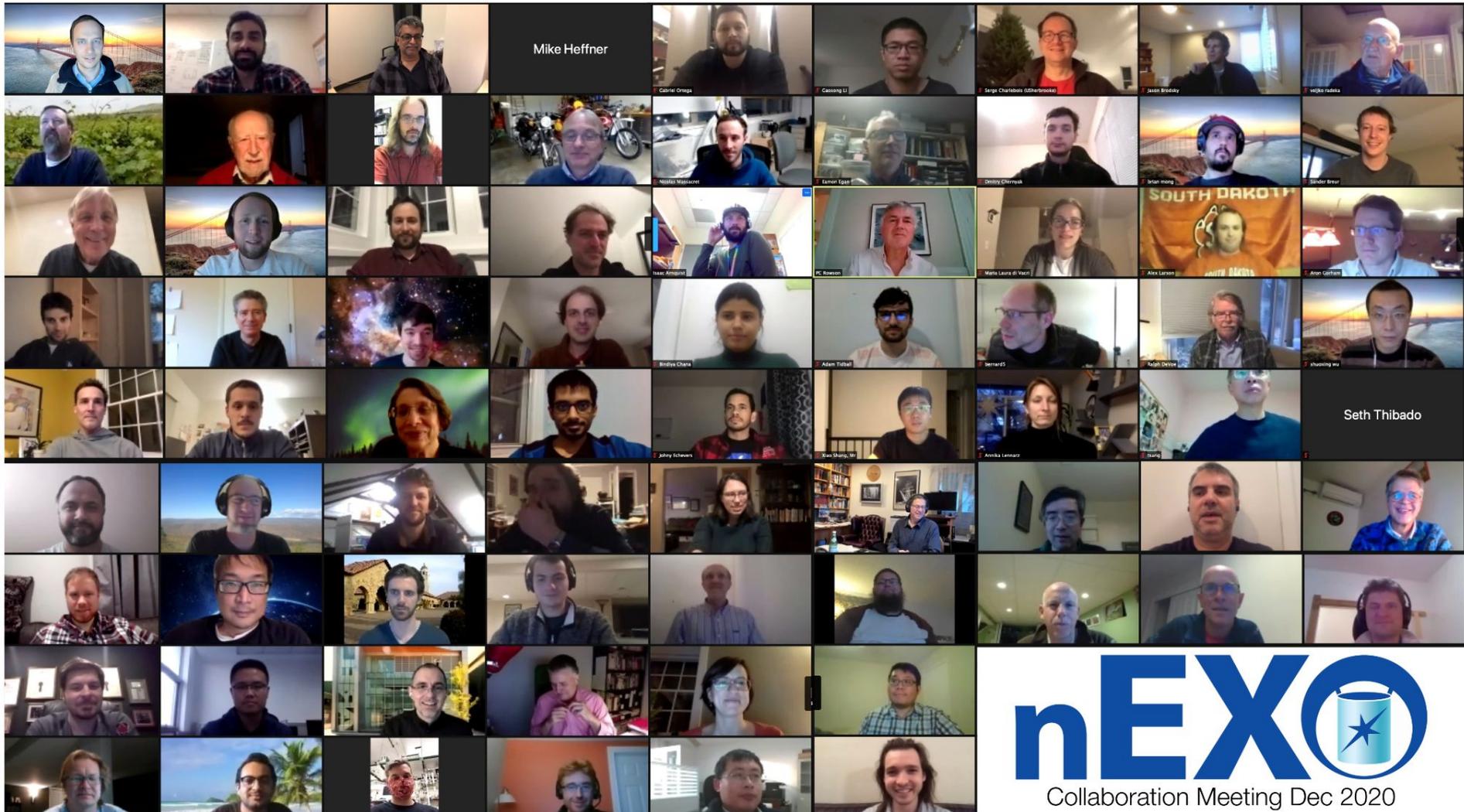
- Sensitivity to Majorana neutrino masses of  $m_{\beta\beta} = 4.7 - 20.3$  meV
  - Matrix elements are dominant systematic uncertainty
- Projections use high-fidelity simulation and data-driven modeling of detector performance

For further details, see [arXiv:2106.16243](https://arxiv.org/abs/2106.16243)

## Other nEXO talks at LIDINE!

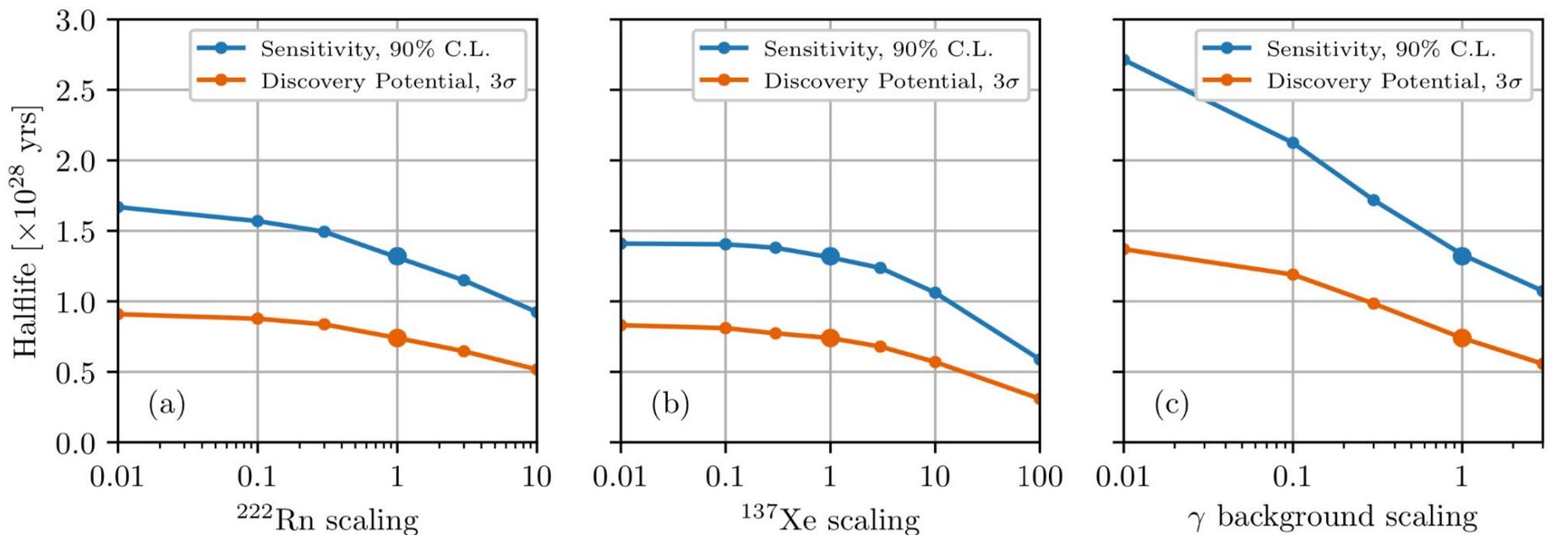
- J. Dalmasson, 12:25pm PT today
- C. Hardy, 2:15pm PT Wednesday
- E. Angelico, 11:00am PT Thursday
- A. Jamil, 10:30am PT Friday

# Thank you!



# Backup

# Dependence of sensitivity on backgrounds



## Depends on $^{222}\text{Rn}$ control

- Dust/materials control is critical
- Online removal under investigation (distillation demonstrated by XENON)

## Depends on cosmogenic muons

- Veto efficiency (70%) supported by detailed simulations, paper soon
- Overburden at SNOLAB (assumed site) is near optimal

## Depends on materials

- $^{222}\text{Rn}$ -related backgrounds are included here
- Significant improvements expected if Ba-tagging is possible in future